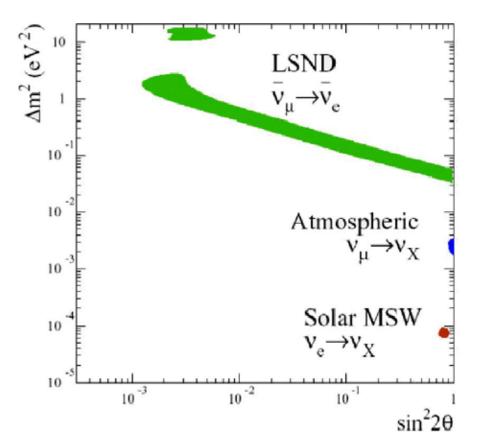
MiniBooNE and Sterile Neutrinos

M. Shaevitz Columbia University WIN 05 Workshop

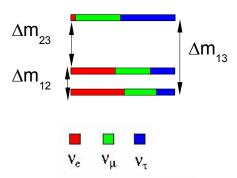
- Extensions to the Neutrino Standard Model: Sterile Neutrinos
- MiniBooNE: Status and Prospects
- Future Directions if MiniBooNE Sees Oscillations

Three Signal Regions



$$P(\nu_{\alpha} \to \nu_{\alpha}) = 1 - \sin^2 2\theta \sin^2 (1.27\Delta m^2 L/E)$$

- LSND $\Delta m^2 = 0.1 10 \text{ eV}^2$, small mixing
- Atmospheric $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$, large mixing
- Solar $\Delta m^2 = 8.0 \times 10^{-5} \text{ eV}^2$, large mixing
- Three distinct neutrino oscillation signals, with $\Delta m_{solar}^2 + \Delta m_{atm}^2 \neq \Delta m_{LSND}^2$
- For three neutrinos, expect $\Delta m_{21}^2 + \Delta m_{32}^2 = \Delta m_{31}^2$

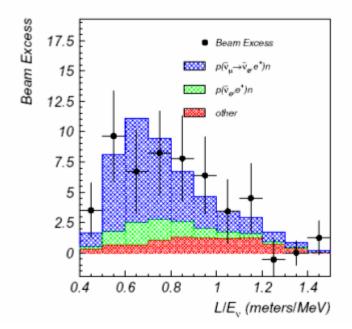


How Can There Be Three Distinct Δm^2 ?

- One of the experimental measurements is wrong
 - Many checks but need MiniBooNE to address LSND
- One of the experimental measurements is not neutrino oscillations
 - Neutrino decay ⇒ Restriction from global fits
 - Neutrino production from flavor violating decays ⇒ Karmen restricts
- Additional "sterile" neutrinos involved in oscillations
 - Still a possibility but probably need (3+2) model
- CPT violation (or CP viol. and sterile v's) allows different mixing for v's and \overline{v} 's
 - Some possibilities still open

LSND Result

- Excess of candidate $\bar{\nu}_e$ events 87.9 \pm 22.4 \pm 6.0 events (3.8 σ) P($\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$) = 0.264 \pm 0.081 %
- Backgrounds in green, red
- Fit to oscillation hypothesis in blue

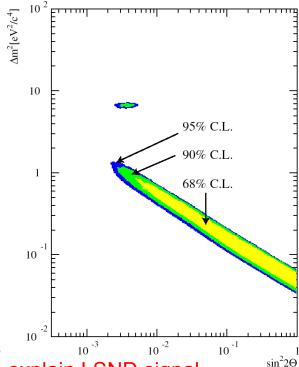


Also Karmen Experiment

 Similar beam and detector to LSND

Closer distance and less target mass ⇒ x10 less sensitive than LSND

 Joint LSND/Karmen analysis gives restricted region (Church et al. hep-ex/0203023)



Also, from Karmen exp.

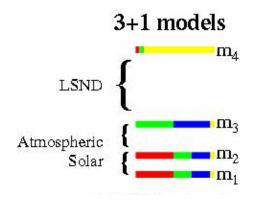
 $\mu^+ \rightarrow e^+ \stackrel{-}{\nu}_e \nu$ unlikely to explain LSND signal

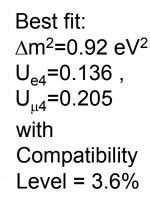
Experimental Situation: Fits of 3+1 and 3+2 Models to Data

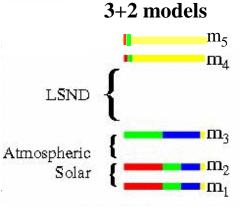
• Global Fits to high Δm^2 oscillations for Short-Baseline exps including LSND positive signal. (M.Sorel, J.Conrad, M.S., hep-ph/0305255)

Is LSND consistent with the upper limits on active to sterile mixing derived from the null short-baseline experiments?

Channel	Experiment	Lowest Δm^2	$\sin^2 2\theta$ Constraint (90% CL)	
	÷n	Reach (90% CL)	High Δm^2	Optimal Δm^2
$\nu_{\mu} ightarrow \nu_{\mathbf{e}}$	LSND	$3 \cdot 10^{-2}$	$> 2.5 \cdot 10^{-3}$	$> 1.2 \cdot 10^{-3}$
	KARMEN	$6 \cdot 10^{-2}$	$< 1.7 \cdot 10^{-3}$	$< 1.0 \cdot 10^{-3}$
	NOMAD	$4 \cdot 10^{-1}$	$< 1.4 \cdot 10^{-3}$	$< 1.0 \cdot 10^{-3}$
$\nu_{\mathbf{e}} \rightarrow \nu_{\mathbf{e}'}$	Bugey	$1 \cdot 10^{-2}$	$< 1.4 \cdot 10^{-1}$	$< 1.3 \cdot 10^{-2}$
	CHOOZ	$7 \cdot 10^{-4}$	$< 1.0 \cdot 10^{-1}$	$< 5 \cdot 10^{-2}$
$\nu_{\mu} \rightarrow \nu_{\mu}$	CCFR84	$6 \cdot 10^{0}$	none	$< 2 \cdot 10^{-1}$
	CDHS	$3 \cdot 10^{-1}$	none	$< 5.3 \cdot 10^{-1}$
$\nu_{\mu} \rightarrow \nu_{\tau}$	NOMAD	$7 \cdot 10^{-1}$	$< 3.3 \cdot 10^{-4}$	$< 2.5 \cdot 10^{-4}$
100	CHORUS	$5 \cdot 10^{-1}$	$< 6.8 \cdot 10^{-4}$	$< 4.5 \cdot 10^{-4}$
$\nu_{\mathbf{e}} \rightarrow \nu_{\tau}$	NOMAD	$6 \cdot 10^{0}$	$< 1.5 \cdot 10^{-2}$	$< 1.1 \cdot 10^{-2}$
	CHORUS	$7 \cdot 10^{0}$	$< 5.1 \cdot 10^{-2}$	$< 4 \cdot 10^{-2}$





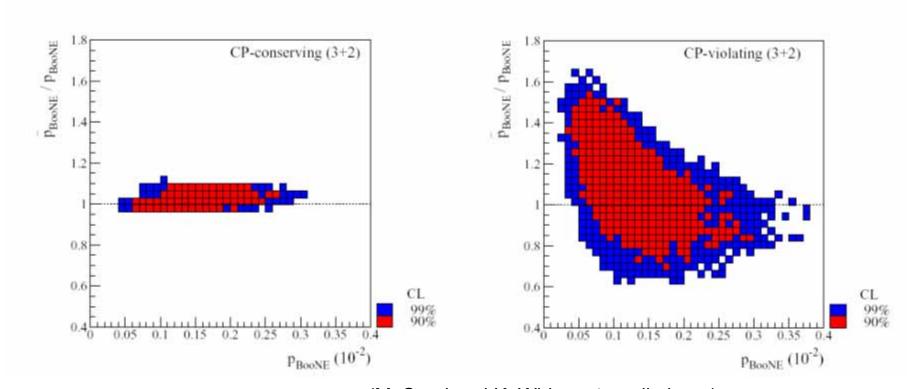


Best Fit: $\Delta m_{41}^2 = 0.92 \text{ eV}^2$ $U_{e4} = 0.121$, $U_{\mu 4} = 0.204$ $\Delta m_{51}^2 = 22 \text{ eV}^2$ $U_{e5} = 0.036$, $U_{\mu 4} = 0.224$ with Compatibility Level = 30%

CP Violation Effect for MiniBooNE in 3+2 Models

- ullet CP-violation is possible when more than one Δm^2 participates in the oscillation
- Compare oscillation probabilities in ν and $\bar{\nu}$ running mode:

$$p_{\rm BooNE} \equiv \langle P(\stackrel{(-)}{\nu_{\mu}} \rightarrow \stackrel{(-)}{\nu_{e}}) \rangle_{\nu \ \rm mode}, \quad \ \bar{p}_{\rm BooNE} \equiv \langle P(\stackrel{(-)}{\nu_{\mu}} \rightarrow \stackrel{(-)}{\nu_{e}}) \rangle_{\bar{\nu} \ \rm mode}$$



(M. Sorel and K. Whisnant, preliminary)

Next Step Is MiniBooNE

 MiniBooNE will be one of the first experiments to check these sterile neutrino models

Investigate LSND Anomaly

Investigate oscillations to sterile neutrino using v_{μ} disappearance

3:1 slope

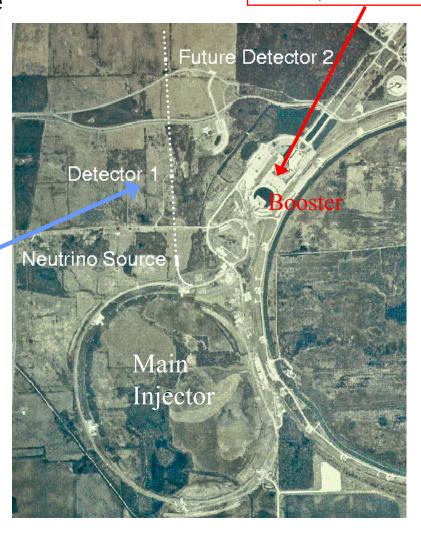
50 ft

24 ft

12 %

40 ft

Use protons from the 8 GeV booster \Rightarrow Neutrino Beam $\langle E_v \rangle \sim 1 \text{ GeV}$



MiniBooNE Collaboration



MiniBooNE consists of about 70 scientists from 13 institutions.

Y. Liu, I. Stancu Alabama

S. Koutsoliotas Bucknell

E. Hawker, R.A. Johnson, J.L. Raaf Cincinnati

T. Hart, R.H. Nelson, E.D. Zimmerman Colorado

A. Aguilar-Arevalo, L.Bugel, L. Coney, J.M. Conrad,

Z. Djurcic, J. Link, J. Monroe, K. McConnel,

D. Schmitz, M.H. Shaevitz, M. Sorel,

G.P. Zeller Columbia

D. Smith Embry Riddle

L.Bartoszek, C. Bhat, S J. Brice, B.C. Brown,

D.A. Finley, R. Ford, F.G.Garcia,

P. Kasper, T. Kobilarcik, I. Kourbanis,

A. Malensek, W. Marsh, P. Martin, F. Mills,

C. Moore, P. Nienaber, E. Prebys,

A.D. Russell, P. Spentzouris, R. Stefanski,

T. Williams Fermilab

D. C. Cox, A. Green, H.-O. Meyer, R. Tayloe *Indiana*

G.T. Garvey, C. Green, W.C. Louis, G.McGregor,

S.McKenney, G.B. Mills, H. Ray, V. Sandberg,

B. Sapp, R. Schirato, R. Van de Water,

D.H. White Los Alamos

R. Imlay, W. Metcalf, M. Sung, M.O. Wascko Louisiana State

J. Cao, Y. Liu, B.P. Roe, H. Yang *Michigan*

A.O. Bazarko, P.D. Meyers, R.B. Patterson,

F.C. Shoemaker, H.A. Tanaka *Princeton*

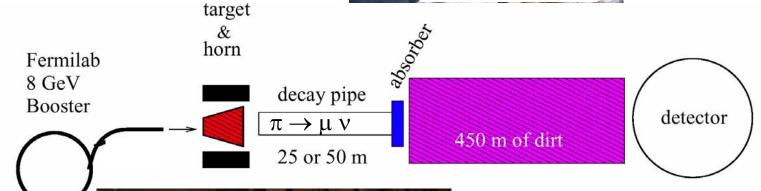
B.T. Fleming Yale

8 GeV Proton Beam Transport

MiniBooNE Neutrino Beam

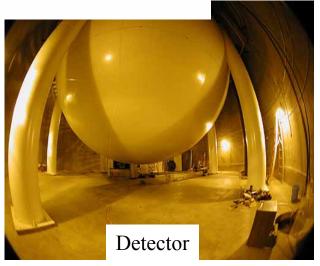


Variable decay pipe length (2 absorbers @ 50m and 25m)



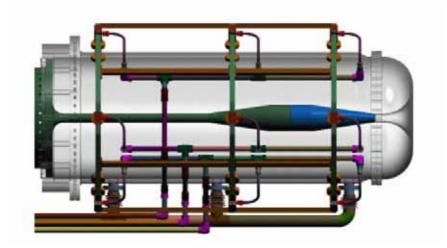
One magnetic Horn, with Be target

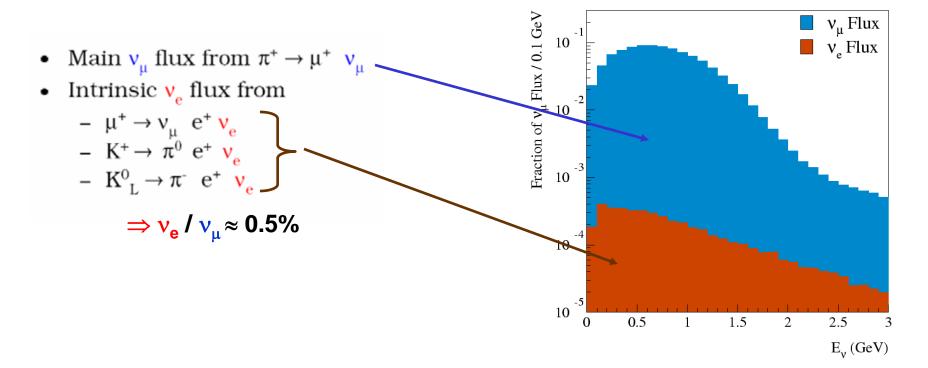




MiniBooNE Horn

- 8 GeV protons impinge on 71cm Be target
- Horn focuses secondaries and increases flux by factor of ~5
- 170 kA pulses, 143 µs long at 5 Hz





The MiniBooNE Detector

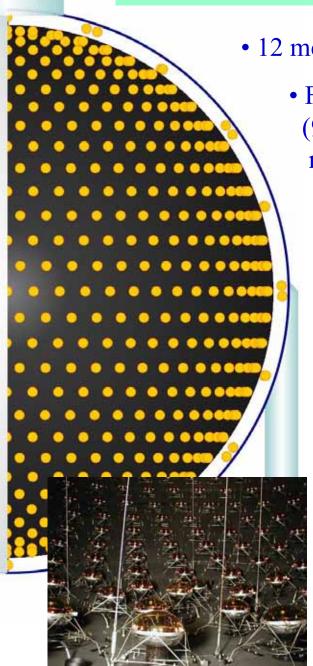




• Filled with 950,000 liters (900 tons) of very pure mineral oil

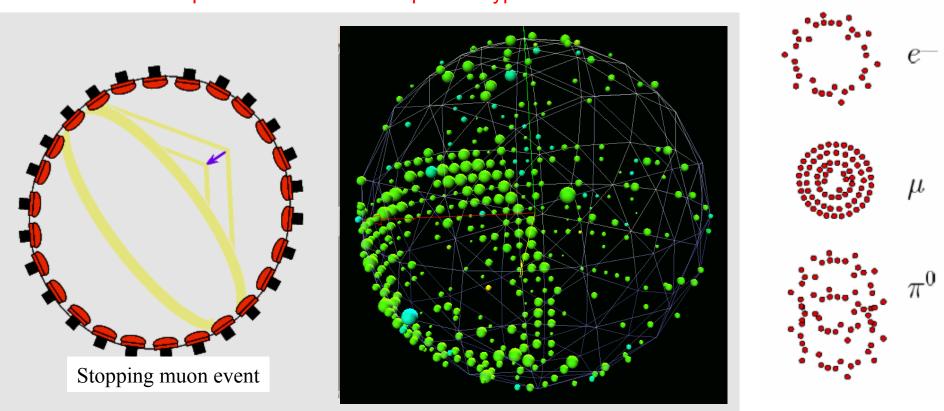
- Light tight inner region with 1280 photomultiplier tubes
- Outer veto region with 241 PMTs.
- Oscillation Search Method:

Look for v_e events in a pure v_μ beam



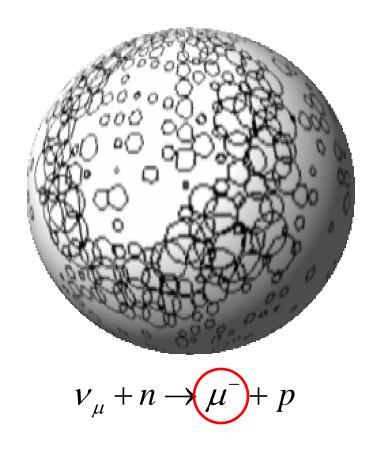
Particle Identification

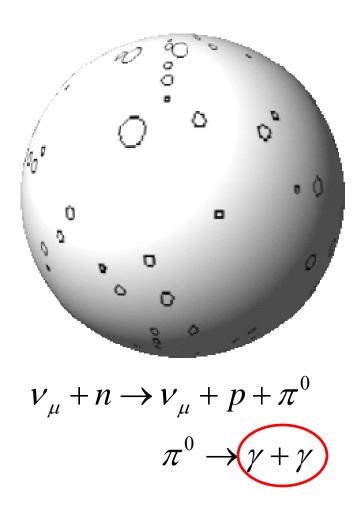
- Separation of ν_{μ} from ν_{e} events
 - Exiting ν_{μ} events fire the veto
 - Stopping $\nu_{\mathfrak{u}}$ events have a Michel electron after a few μsec
 - Also, scintillation light with longer time constant ⇒ enhanced for slow pions and protons
 - Čerenkov rings from outgoing particles
 - Shows up as a ring of hits in the phototubes mounted inside the MiniBooNE sphere
 - Pattern of phototube hits tells the particle type



Example Cerenkov Rings

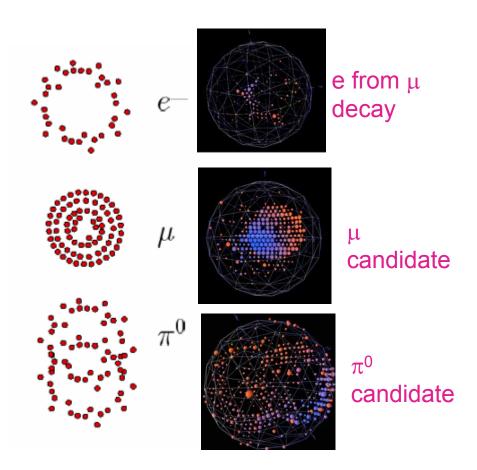
Size of circle is proportional to the light hitting the photomultiplier tube





Particle ID Algorithms

- Identify events using hit topology
- Use a "boosted tree" algorithm to separate e, mu, pi, delta
 - More stable than neural network in performance and less sensitivity to MC optical model (See B. Roe et al. NIM A543 (2005))
- PID Vars
 - Reconstructed physical observables
 - Track length, particle production angle relative to beam direction
 - Auxiliary quantities
 - Timing, charge related : early/prompt/late hit fractions, charge likelihood
 - Geometric quantities
 - Distance to wall



Neutrino events

beam comes in spills @ up to 5 Hz each spill lasts 1.6 μsec

trigger on signal from Booster read out for 19.2 μsec

no high level analysis needed to see neutrino events

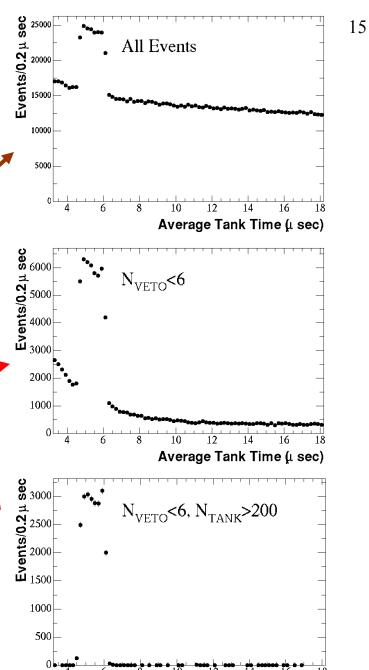
backgrounds: cosmic muons $\leftarrow N_{Veto} < 6$ Cut decay electrons ← N_{Tank}<200 Cut

simple cuts reduce non-beam backgrounds to ~10⁻³

v event every 1.5 minutes

Current Collected data:

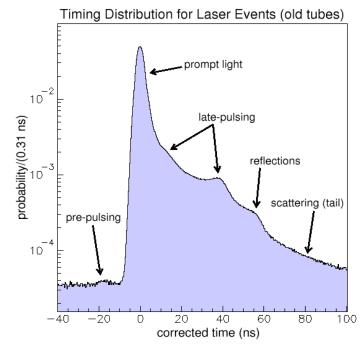
~600k neutrino candidates for 5.6×10^{20} protons on target



Average Tank Time (a sec)

Optical Model

- Light Creation
 - Cerenkov well known
 - Scintillation
 - yield
 - spectrum
 - decay times
- Light Propagation
 - Fluoresence
 - rate
 - spectrum
 - decay times
 - Scattering
 - Rayleigh
 - Particulate (Mie)
 - Absorption



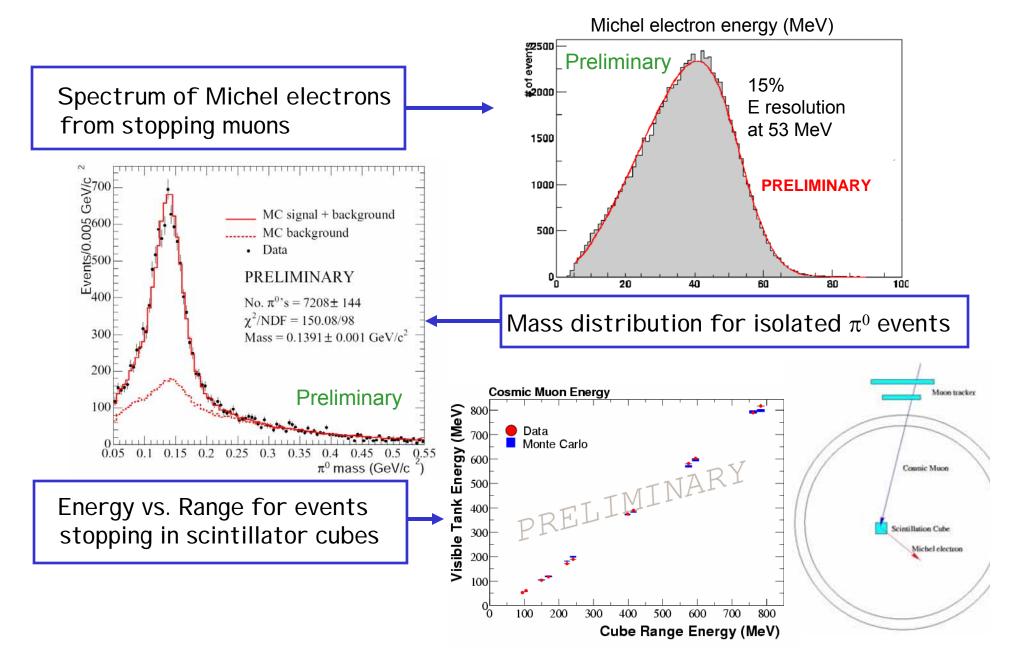
In Situ

Cosmics muons, Michel electrons, Laser

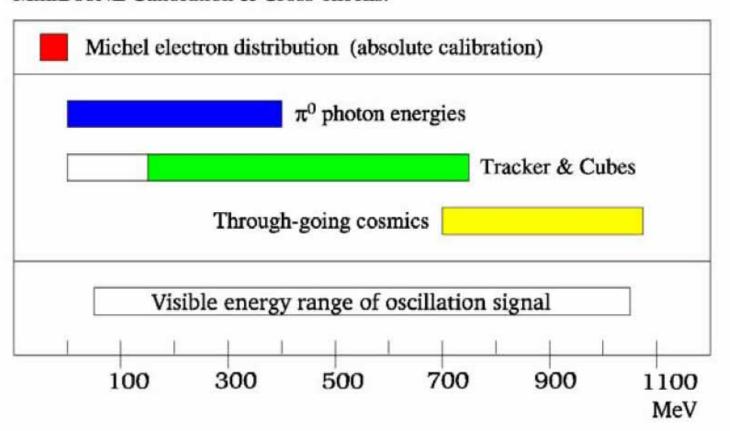
Ex Situ

- Scintillation from p beam (IUCF)
- Scintillation from cosmic μ (Cincinnati)
- Fluorescence Spectroscopy (FNAL)
- Time resolved spectroscopy (JHU)
- Attenuation (Cincinnati)

Energy Calibration Signals

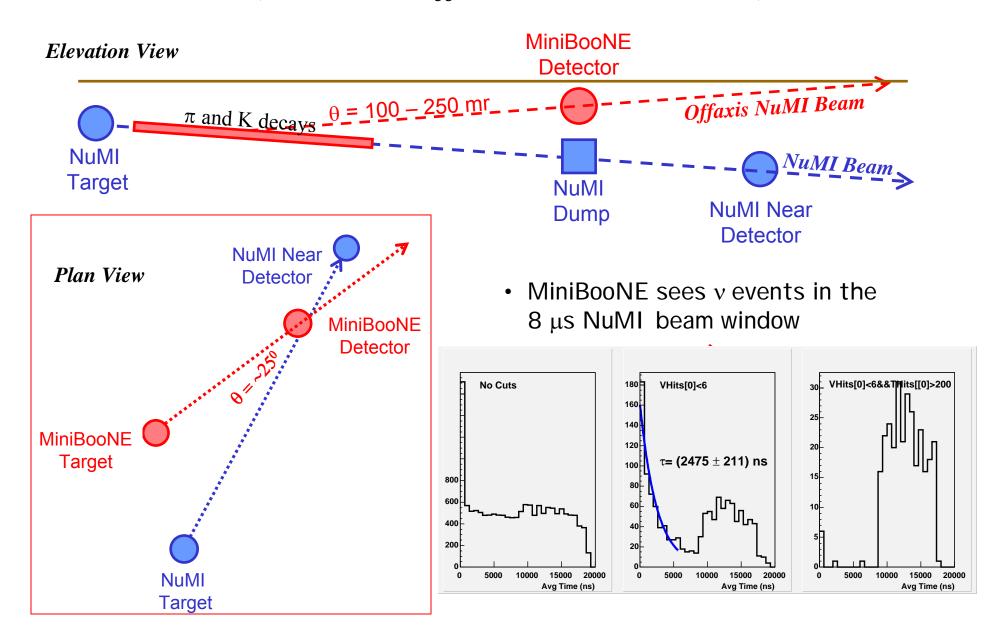


MiniBooNE Calibration & Cross-checks:



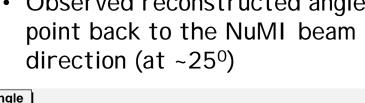
NuMI Beam Events in MiniBooNE

(World's 1st Offaxis Neutrino Beam !!)

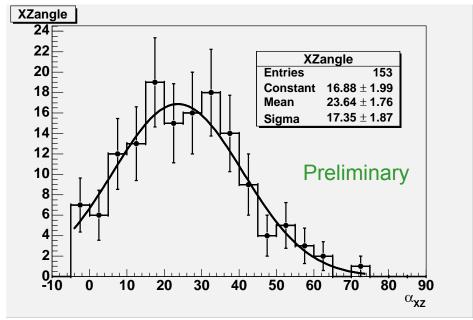


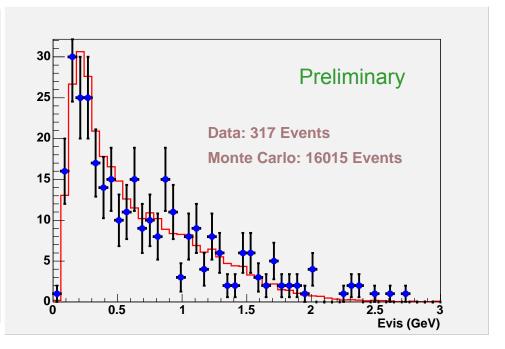
NuMI Offaxis Events Agree with Monte Carlo Prediction

 Observed reconstructed angle point back to the NuMI beam direction (at ~25°)



 Data to Monte Carlo comparison of reconstructed E_{visible} for contained events





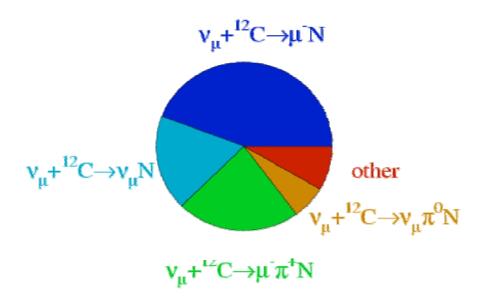
⇒ NuMI Offaxis beam will be a calibration beam for MiniBooNE (and we can look at electron neutrino interactions)

Oscillation Analysis: Status and Plans

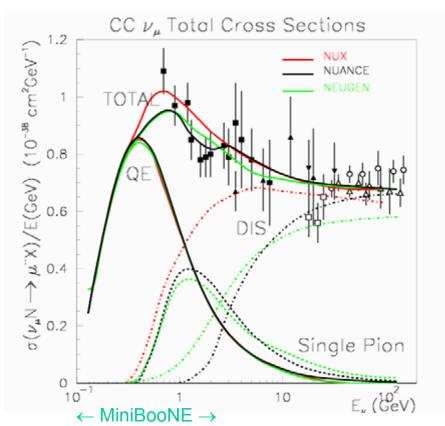
- Blind (or "Closed Box") ν_e appearance analysis you can see all of the info on some events or some of the info on all events but you cannot see all of the info on all of the events
- Other analysis topics give early interesting physics results and serve as a cross check and calibration before "opening the ν_e box"
 - Cross section measurements for low-energy v processes
 - v_{μ} disappearance oscillation search
 - Studies of ν_{μ} NC π^0 production \Rightarrow coherent (nucleus) vs nucleon
 - Studies of v_{μ} NC elastic scattering \Rightarrow Measurements of Δs (strange quark spin contribution)

Low Energy Neutrino Cross sections

MiniBooNE Events Fractions



 MiniBooNE will measure the cross sections for all of these processes

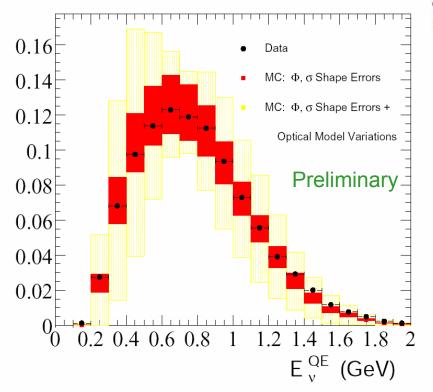


On the Road to a ν_{μ} Disappearance Result

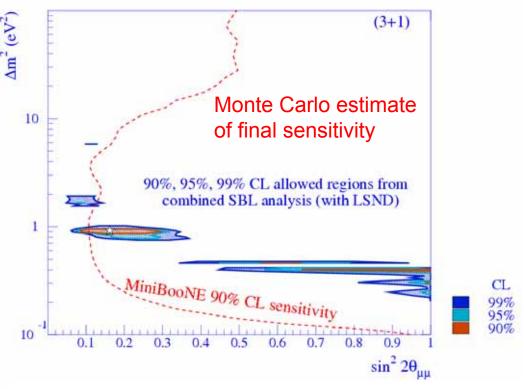
• Use ν_{μ} quasi-elastic events

$$\nu_{\mu}$$
+n \rightarrow μ^- +p

- Events can be isolated using single ring topology and hit timing
- Excellent energy resolution
- High statistics: ~30,000 events now (Full sample: ~500,000)



- E_{ν} distribution well understood from pion production by 8 GeV protons
 - Sensitivity to $\nu_{\mu} \rightarrow \nu_{\mu}$ disappearance oscillations through shape of E_{ν} distribution

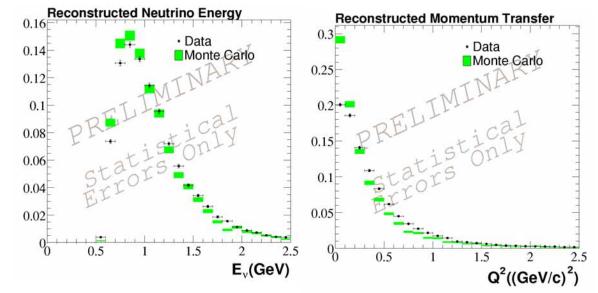


Will be able to cover a large portion of 3+1 models

Neutrino Single Pion Production Cross Sections

• Charged current π^+ events Resonant

$$\begin{array}{c} \nu_{\mu}+\mathbf{p}\rightarrow\mu^{-}+\Delta^{++}\\ \longrightarrow\mathbf{p}+\pi^{+}\\ \text{Coherent}\\ \nu_{\mu}+\mathbf{N}\rightarrow\mu^{-}+\mathbf{N}+\pi^{+} \end{array}$$

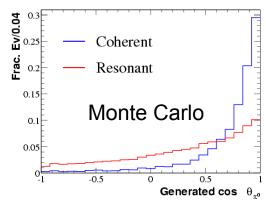


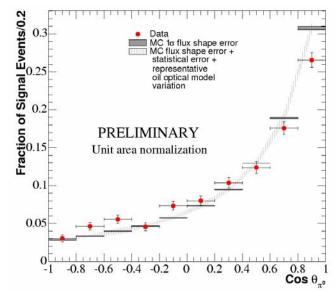
• Neutral current π^0 events Resonant

$$\nu_{\mu} + \mathbf{n} \rightarrow \nu_{\mu} + \Delta^{0} \longrightarrow \mathbf{n} + \pi^{0}$$

Coherent

$$\nu_{\mu} \text{+} \textbf{N} \rightarrow \nu_{\mu} \text{+} \textbf{N} + \pi 0$$

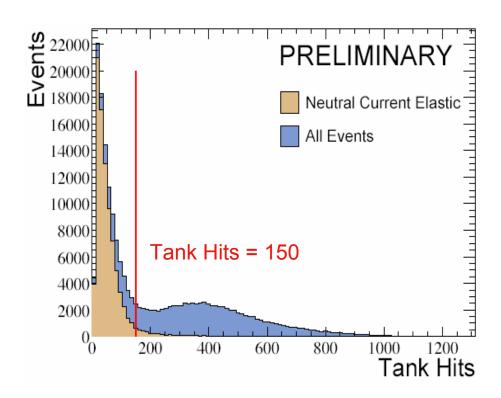




Investigations of ν_{μ} NC elastic scattering

$$V_{\mu} + p \rightarrow V_{\mu} + p$$

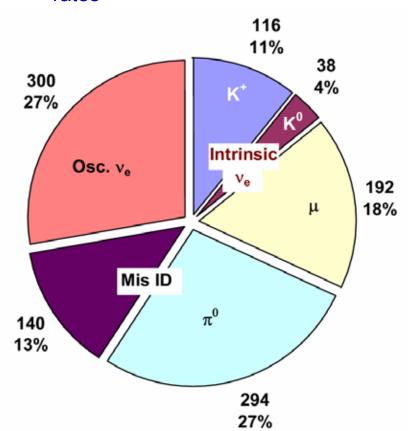
- Study scint. properties of oil, low E response of detector
 - Reconstruct p energy from scint. light
- Measure $\sigma (v_{\mu} + p \rightarrow v_{\mu} + p)$
 - Help understand scint. light for v_e osc analysis
- σ(NCE) / σ (CCQE)
 - Measure ∆s (component of proton spin carried by strange quarks)



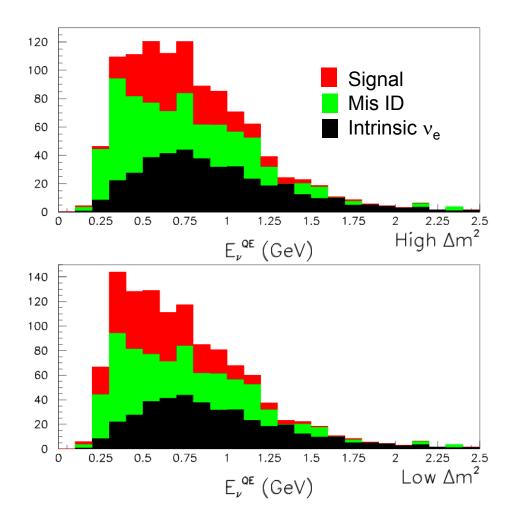
Tank hits < 150, veto < 6, 1 sub-event : ε = 70%, purity = 80%

Estimates for the $\nu_{\mu} \rightarrow \nu_{e}$ Appearance Search

- Look for appearance of ν_{e} events above background expectation
 - Use data measurements both internal and external to constrain background rates

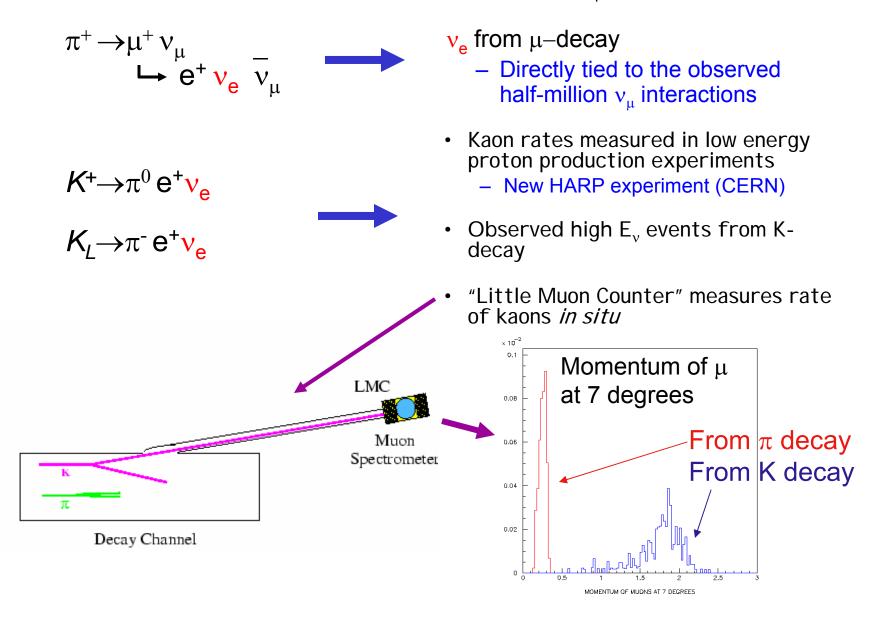


• Fit to E_{ν} distribution used to separate background from signal.



Intrinsic v_e in the beam

Small intrinsic v_e rate \Rightarrow Event Ratio v_e/v_μ =6x10⁻³

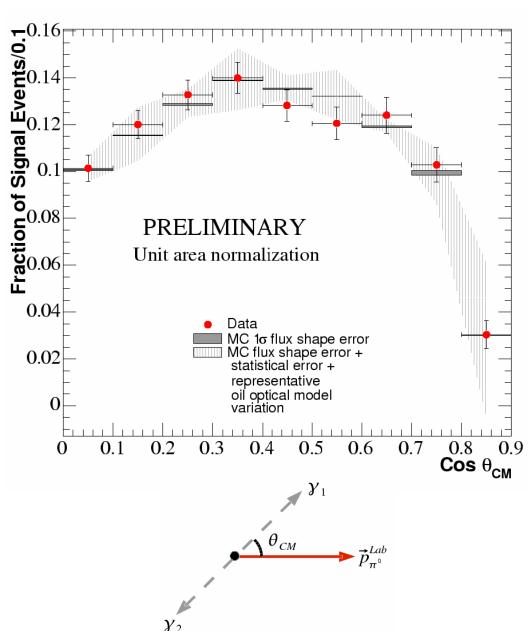


Mis-identification Backgrounds

• Background mainly from NC π^0 production

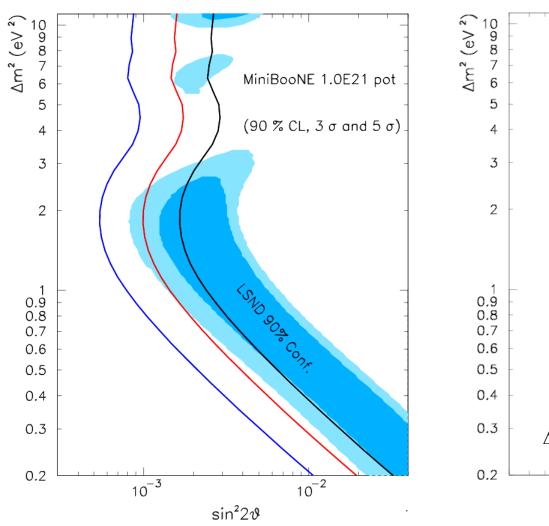
```
v_{\mu} + p \rightarrow v_{\mu} + p + \pi^{0} followed by \pi^{0} \rightarrow \gamma \gamma where one \gamma is lost because it has too low energy or have overlapping rings
```

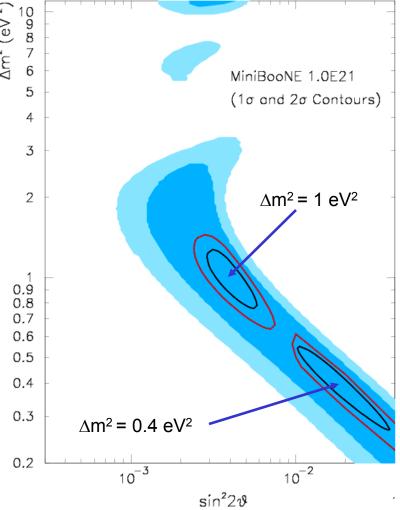
- Over 99.5% of these events are identified and the π^0 kinematics are measured
 - ⇒ Can constrain this background directly from the observed data



MiniBooNE Oscillation Sensitivity

- Oscillation sensitivity and measurement capability
 - Data sample corresponding to 1x10²¹ pot
 - Systematic errors on the backgrounds average ~5%

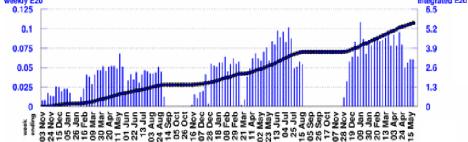




• In its 30 year history, the Fermilab Booster has never worked this hard and this well

Poforo NuMI turn on word averaging

- Before NuMI turn-on were averaging ...
 7x10¹⁶ protons/hour
- Co-running with NuMI now averages ...
 3.5x10¹⁶ protons/hour



Have now reached 5.6×10^{20} protons on target in total

- Already have world's largest v dataset in the 1 GeV region
- Physics results show that reconstruction and analysis algorithms are working well
- Plan is to "open the ν_e appearance box" when the analysis has been substantiated and when sufficient data has been collected for a definitive result
 - ⇒ Estimate is before the end of 2005
- Which then leads to the question of the next step
 - If MiniBooNE sees no indications of oscillations with ν_μ
 - \Rightarrow Need to run with $\stackrel{-}{\nu}_{\!_{1}}$ since LSND signal was $\stackrel{-}{\nu}_{\!_{1}} \!\!\! \rightarrow \stackrel{-}{\nu}_{\!_{2}}$
 - If MiniBooNE sees an oscillation signal⇒ Then

Experimental Program with Sterile Neutrinos

If sterile neutrinos then many mixing angles, CP phases, and Δm^2 to include

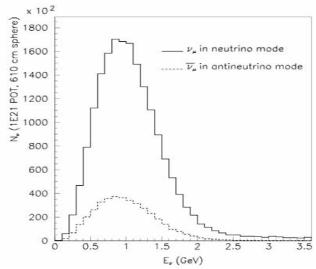
• Measure number of extra masses Δm_{14}^2 , Δm_{15}^2 ...

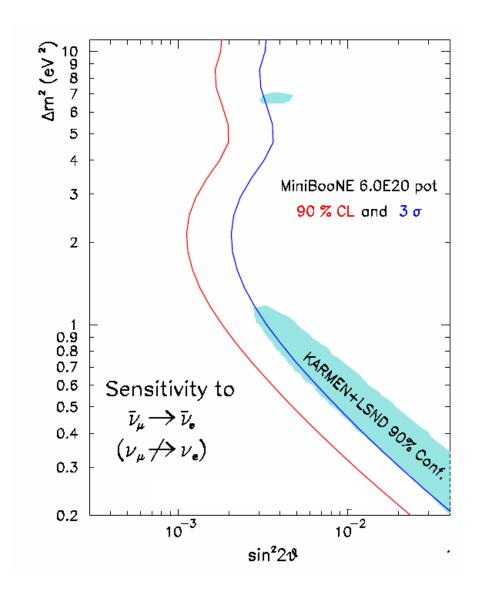
Measure mixings
 Could be many small angles

 Oscillations to sterile neutrinos could effect long-baseline measurements and strategy • Compare $\nu_{_{\!\!\!\! L}}$ and $\, \overline{\nu}_{_{\!\!\! L}}$ oscillations \Rightarrow CP and CPT violations

If MiniBooNE sees $\nu_{\mu} \rightarrow \nu_{e}$ (or not) then: _ Run BooNE with anti-neutrinos for $\nu_{\mu} \rightarrow \nu_{e}$

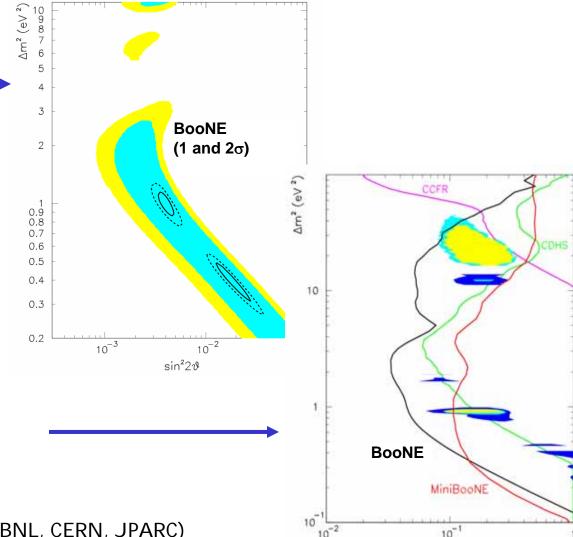
- Direct comparison with LSND
- Are v_{μ} and \overline{v}_{μ} the same?
 - Mixing angles, Δm^2 values
- Explore CP (or CPT) violation by comparing v_{μ} and v_{μ} results
- Running with antineutrinos takes about x2 longer to obtain similar sensitivity





Next Step: BooNE: Two Detector Exp.

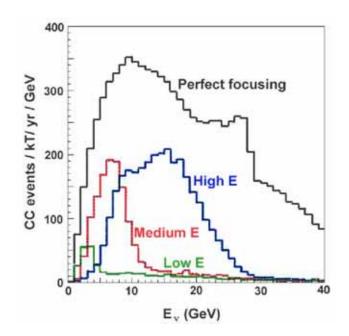
- Precision measurement of oscillation parameters
 - $-\sin^2 2\theta$ and Δm^2
 - Map out the nxn mixing matrix
- Determine how many high mass Δm² 's
 - 3+1, 3+2, 3+3
- Show the L/E oscillation dependence
 - Oscillations or v decay or ???
- Explore disappearance measurement in high Δm^2 region
 - Probe oscillations to sterile neutrinos



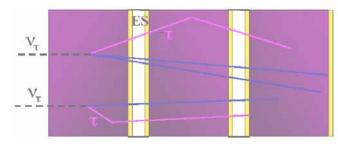
(These exp's could be done at FNAL, BNL, CERN, JPARC)

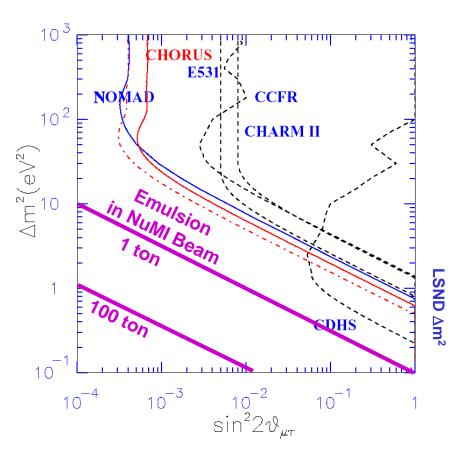
Another Next Step: Do $\nu_{\mu}\!\!\to\!\!\nu_{\tau}$ Appearance Experiment at High Δm^2

- Appearance of v_{τ} would help sort out the mixings through the sterile components
- Need moderately high neutrino energy to get above the 3.5 GeV τ threshold (~6-10 GeV)
- Example: NuMI Med energy beam 8 GeV with detector at L=2km (116m deep)



Emulsion Detector or Liquid Argon





Conclusions

- Neutrinos have been surprising us for some time and will most likely continue to do so
- Although the "neutrino standard model" can be used as a guide,

the future direction for the field is going to be determined by what we discover from experiments.

Sterile neutrinos may open up a whole

v area to explore