

# Neutrino Physics Prospects of Neutrinoless Double-Beta Decay

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WIN 2005  
Delphi, Greece  
June 10, 2005

# Compelling Evidences for $\nu$ -Oscillations

–  $\nu_{\text{atm}}$ : **SK** UP-DOWN ASYMMETRY

$\theta_{23}$ -,  $L/E$ - dependences of  $\mu$ -like events

Dominant  $\nu_{\mu} \rightarrow \nu_{\tau}$  K2K; MINOS, CNGS

–  $\nu_{\odot}$ : Homestake, Kamiokande, SAGE, GALLEX/GNO  
Super-Kamiokande, SNO; KamLAND

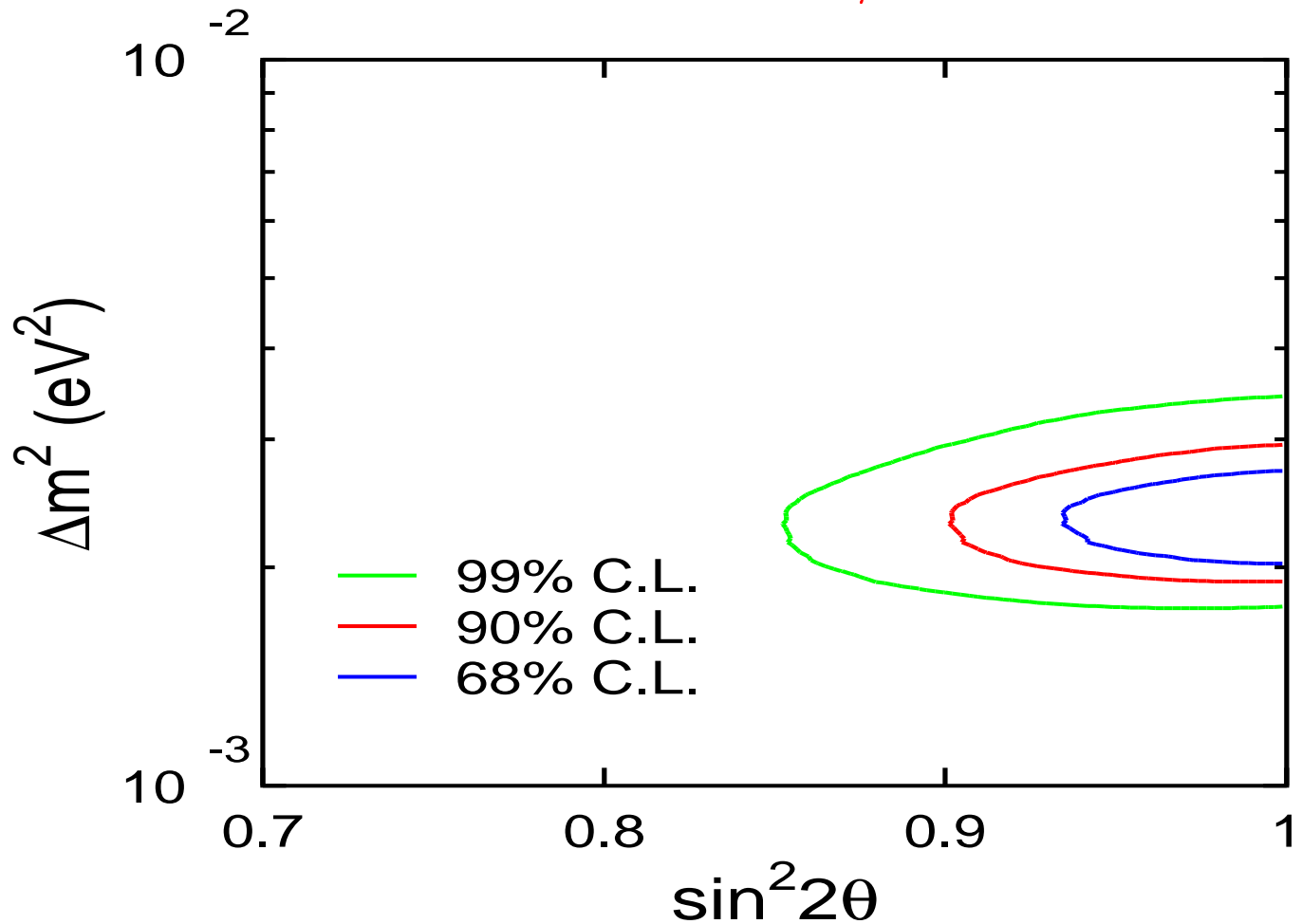
Dominant  $\nu_e \rightarrow \nu_{\mu, \tau}$  BOREXINO, ..., LowNu

– LSND

Dominant  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$  MiniBOONE

$$\nu_{lL} = \sum_{j=1} U_{lj} \nu_{jL} \quad l = e, \mu, \tau. \quad (1)$$

## SK: Atmospheric $\nu$ Data, $L/E$



$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2 = 2.1 \times 10^{-3} \text{ eV}^2, \quad \sin^2 2\theta_{\text{atm}} \equiv \sin^2 2\theta_{23} = 1.0 ;$$

$$\Delta m_{31}^2 = (1.3 - 4.2) \times 10^{-3} \text{ eV}^2, \quad \sin^2 2\theta_{23} \geq 0.85, \quad 99.73\% \text{ C.L.}$$

$$+ \text{K2K} : \quad \Delta m_{31}^2 = (1.4 - 3.3) \times 10^{-3} \text{ eV}^2, \quad 99.73\% \text{ C.L.}$$

M. Maltoni et al., hep-ph/0405172

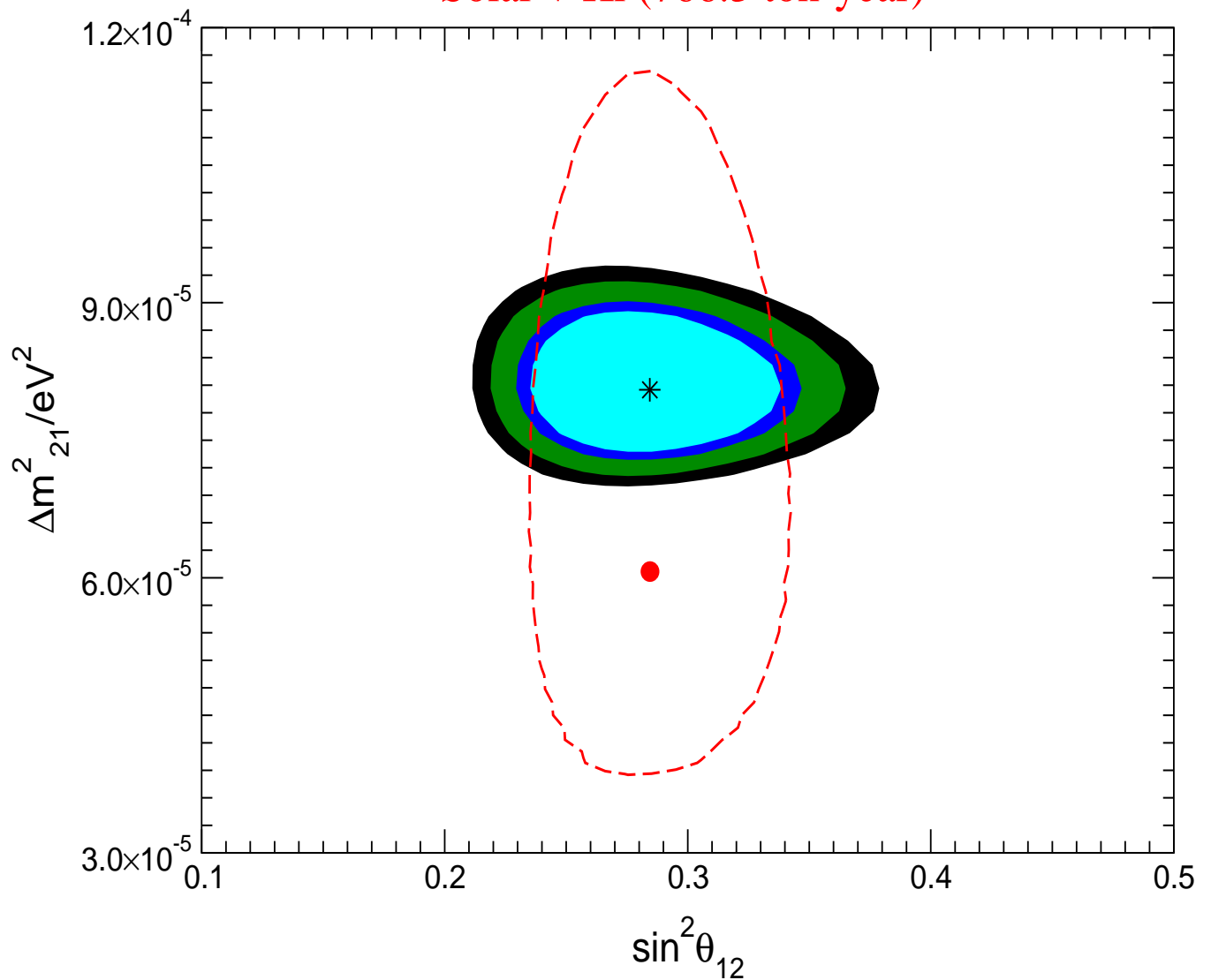
- sign of  $\Delta m_{\text{atm}}^2$  not determined;

$$3\text{-}\nu \text{ mixing: } \Delta m_{31}^2 > 0, \quad m_1 < m_2 < m_3 \text{ (NH);}$$

$$\Delta m_{31}^2 < 0, \quad m_3 < m_1 < m_2 \text{ (IH).}$$

- If  $\theta_{23} \neq \frac{\pi}{4}$ :  $\theta_{23}, (\frac{\pi}{4} - \theta_{23})$  ambiguity.

## Solar + KI (766.3 ton-year)



$$\Delta m_{\odot}^2 \equiv \Delta m_{21}^2 = 8.0 \times 10^{-5} \text{ eV}^2, \quad \sin^2 \theta_{\odot} \equiv \sin^2 \theta_{12} = 0.31 ;$$
$$\cos 2\theta_{12} = 0.39; \quad \cos 2\theta_{12} > 0.28, \quad 95\% \text{ C.L.}$$

- $\sin^2 \theta_{12} = 0.50$  excluded at  $> 6$  s.d.
- High-LMA excluded at  $\sim 3.3$  s.d.

A.Bandyopadhyay, S.Choubey, S.Goswami, S.T.P., D.P.Roy,  
hep-ph/0406328;

M. Maltoni et al., hep-ph/0405172

### 3- $\nu$ Mixing Analysis: $\Delta m_{\odot}^2 \ll |\Delta m_{\text{atm}}^2|$

$$P_{\odot}^{3\nu} \cong \sin^4 \theta_{13} + \cos^4 \theta_{13} P_{\odot}^{2\nu},$$

$$P_{\odot}^{2\nu} = \bar{P}_{\odot}^{2\nu} + P_{\odot}^{2\nu}{}_{\text{osc}},$$

$$\bar{P}_{\odot}^{2\nu} = \frac{1}{2} + \left(\frac{1}{2} - P'\right) \cos 2\theta_{12}^m(t_0) \cos 2\theta_{12} \quad (\theta_{12} \equiv \theta_{\odot}),$$

W. Haxton, S. Parke, 1986

$$N_e \rightarrow N_e \cos^2 \theta_{13},$$

$$P' = \frac{e^{-2\pi r_0 \frac{\Delta m^2}{2E}} \sin^2 \theta - e^{-2\pi r_0 \frac{\Delta m^2}{2E}}}{1 - e^{-2\pi r_0 \frac{\Delta m^2}{2E}}}$$

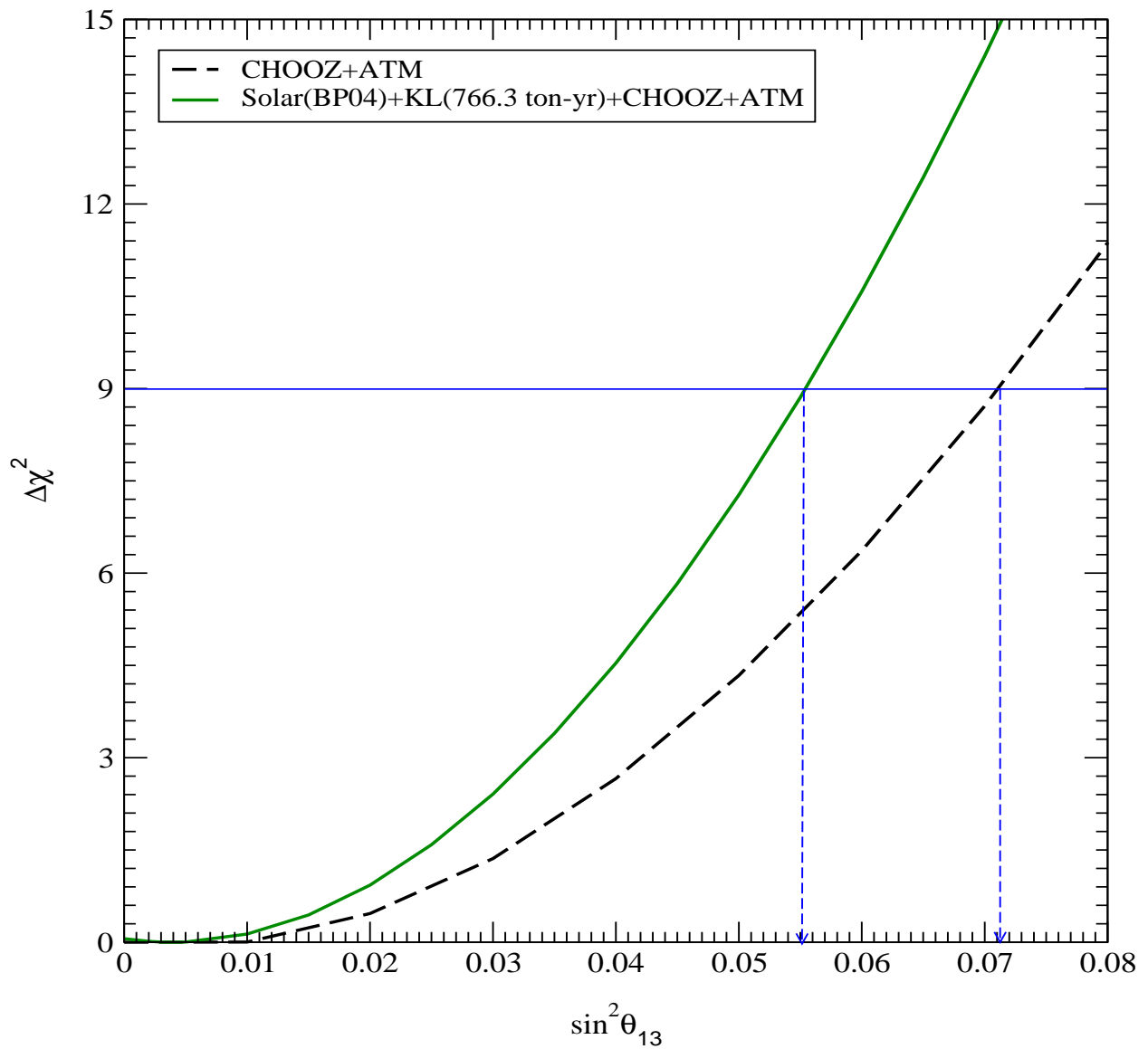
S.T.P., 1988

$$\text{LMA: } P' \ll 1, \quad \langle P_{\odot}^{2\nu}{}_{\text{osc}} \rangle \cong 0$$

J. Rich, S.T.P., 1988

$$P_{\text{KL}}^{3\nu} \cong \sin^4 \theta_{13} + \cos^4 \theta_{13} \left[ 1 - \sin^2 2\theta_{12} \sin^2 \left( \frac{\Delta m_{\odot}^2}{4E} L \right) \right]$$

$$P_{\text{CHOOZ}}^{3\nu} \cong 1 - \sin^2 2\theta_{13} \sin^2 \left( \frac{\Delta m_{\text{atm}}^2}{4E} L \right)$$



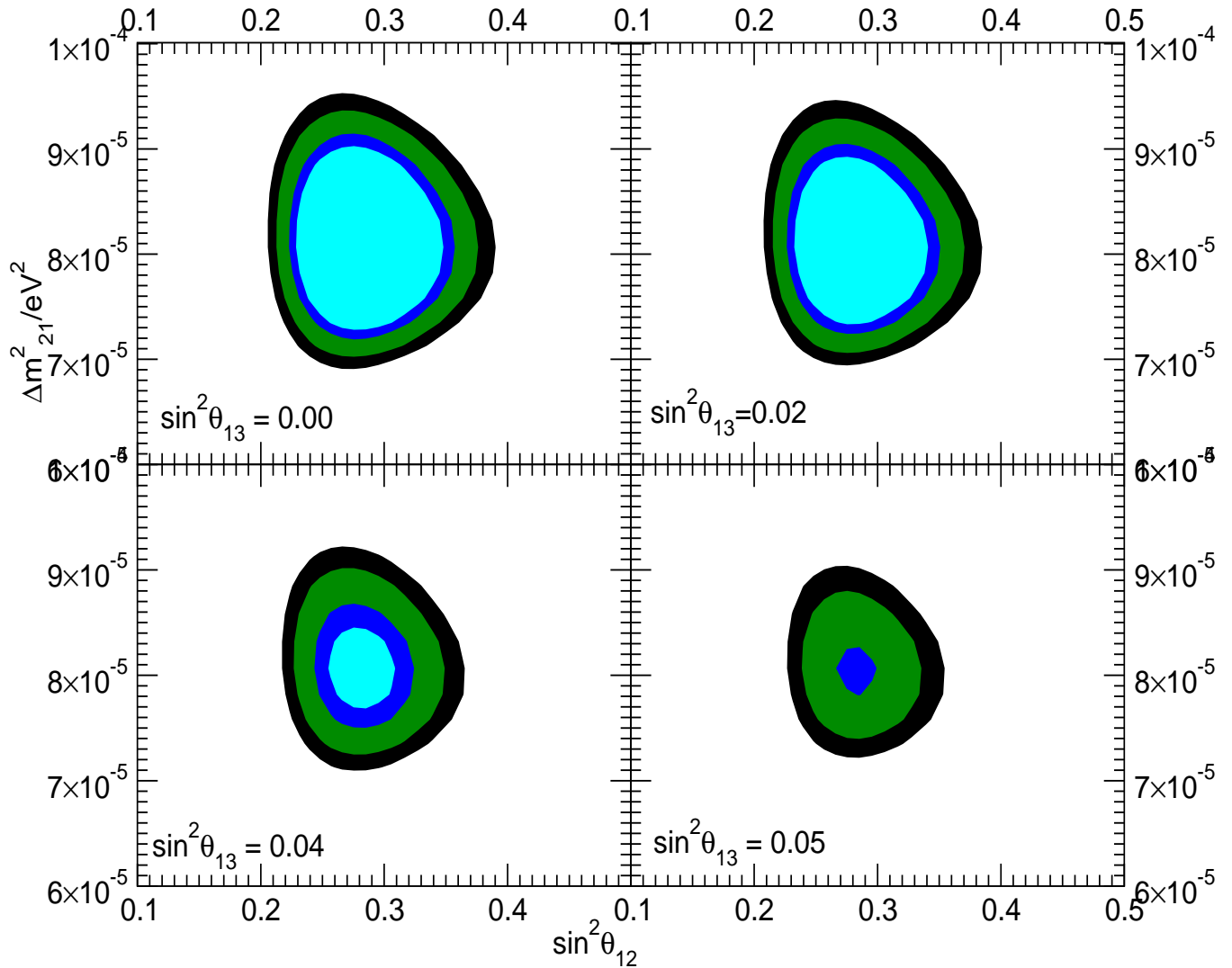
$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2 = (1.3 - 4.2) \times 10^{-3} \text{ eV}^2, \text{ 3 s.d. (SK : } \nu'04)$$

SK: E. Kearns, talk at  $\nu'04$ , June 2004, Paris

- $\sin^2 \theta_{13} < 0.05$  at 99.73% C.L.

A. Bandyopadhyay et al., hep-ph/0406328

Solar + KamLAND(766.3 ton year) + CHOOZ



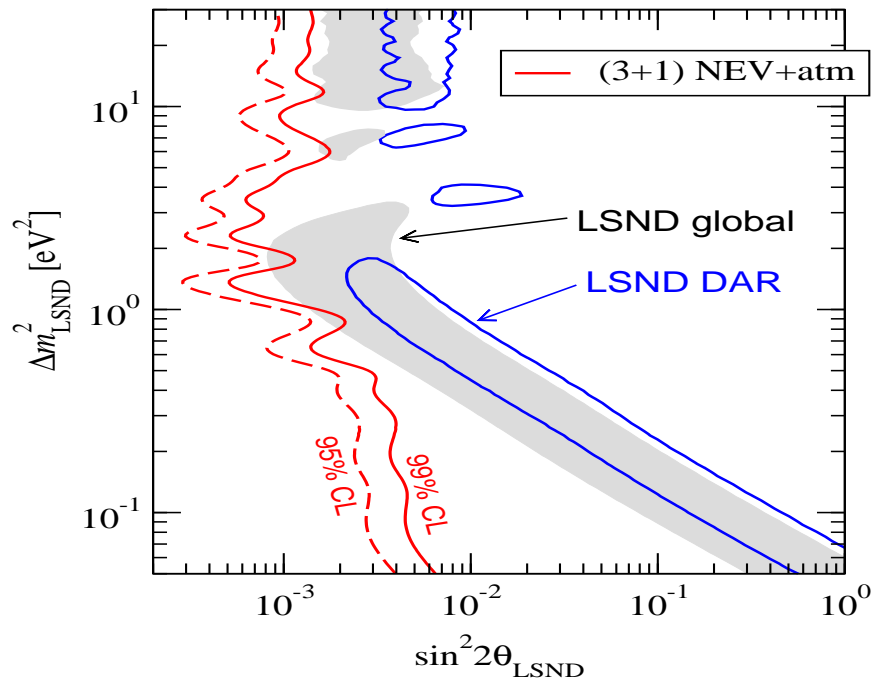
$\sin^2 \theta_{13} = 0$ , 95% C.L. :

$\Delta m^2_{21} = (7.2 - 8.6) \times 10^{-5} \text{ eV}^2$  ,  $\sin^2 \theta_{12} = (0.26 - 0.36)$  ;  
 $\cos 2\theta_{12} = 0.40$ ;  $\cos 2\theta_{12} > 0.28$ , 95% C.L.;

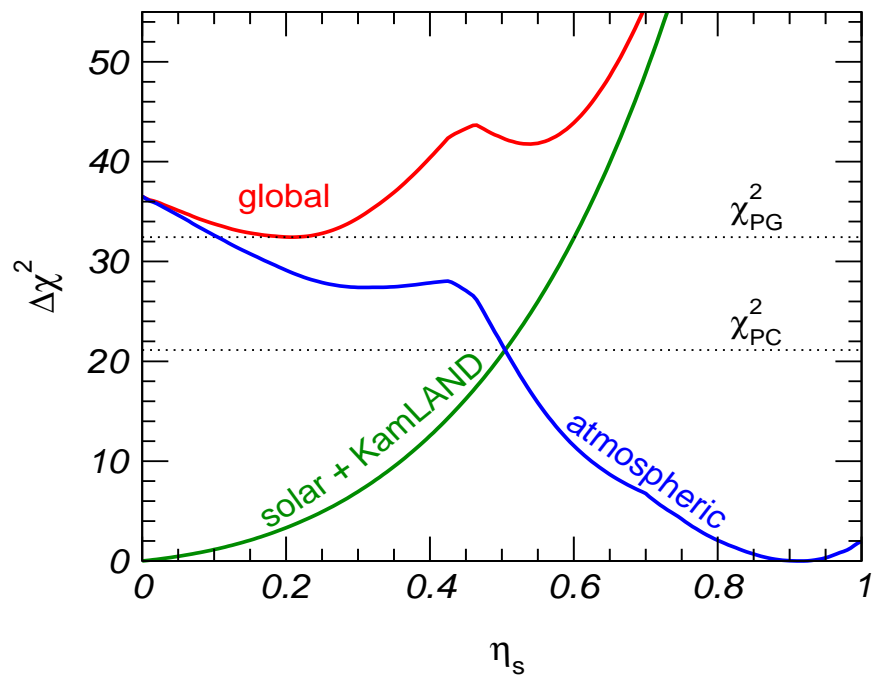
- $\sin^2 \theta_{12} = 0.50$  excluded at  $> 6$  s.d.

A. Bandyopadhyay et al, hep-ph/0406328

(3+1)



(2+2)



M. Maltoni et al., hep-ph/0405172



# Three Neutrino Mixing

$$\nu_{lL} = \sum_{j=1}^3 U_{lj} \nu_{jL} . \quad (2)$$

$U$  is the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) neutrino mixing matrix,

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \quad (3)$$

- $U$  -  $n \times n$  unitary:

$n$	2	3	4
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mixing angles:	$\frac{1}{2}n(n-1)$	1	3	6
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CP-violating phases:

• $\nu_j$ - Dirac:	$\frac{1}{2}(n-1)(n-2)$	0	1	3
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• $\nu_j$ - Majorana:	$\frac{1}{2}n(n-1)$	1	3	6
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$n = 3$ : 1 Dirac and  
2 additional CP-violating phases, Majorana phases

S.M. Bilenky, J. Hosek, S.T.P., 1980;  
J. Schechter, J.W.F. Valle, 1980;  
M. Doi, T. Kotani, E. Takasugi, 1981

# Majorana Neutrinos

Can be defined in QFT using fields or states.

Fields:  $\chi_k(x)$  - 4 component (spin 1/2), complex,  $m_k$

Majorana condition:

$$C (\bar{\chi}_k(x))^T = \xi_k \chi_k(x), \quad |\xi_k|^2 = 1$$

- Invariant under proper Lorentz transformations.
- Reduces by 2 the number of components in  $\chi_k(x)$ .

Implications:

$$U(1) : \chi_k(x) \rightarrow e^{i\alpha} \chi_k(x) - \text{impossible}$$

- $\chi_k(x)$  cannot absorb phases.
- $Q_{U(1)} = 0 : Q_{el} = 0, L_l = 0, L = 0, \dots$
- $\chi_k(x)$ : 2 spin states of a spin 1/2 absolutely neutral particle
- $\chi_k \equiv \bar{\chi}_k$

Propagators:  $\Psi(x)$ -Dirac,  $\chi(x)$ -Majorana

$$\langle 0|T(\Psi_\alpha(x)\bar{\Psi}_\beta(y))|0 \rangle = S_{\alpha\beta}^F(x-y) ,$$

$$\langle 0|T(\Psi_\alpha(x)\Psi_\beta(y))|0 \rangle = 0 , \quad \langle 0|T(\bar{\Psi}_\alpha(x)\bar{\Psi}_\beta(y))|0 \rangle = 0 .$$

$$\langle 0|T(\chi_\alpha(x)\bar{\chi}_\beta(y))|0 \rangle = S_{\alpha\beta}^F(x-y) ,$$

$$\langle 0|T(\chi_\alpha(x)\chi_\beta(y))|0 \rangle = -\xi^* S_{\alpha\kappa}^F(x-y) C_{\kappa\beta} ,$$

$$\langle 0|T(\bar{\chi}_\alpha(x)\bar{\chi}_\beta(y))|0 \rangle = \xi C_{\alpha\kappa}^{-1} S_{\kappa\beta}^F(x-y)$$

$$U_{CP} \chi(x) U_{CP}^{-1} = \eta_{CP} \gamma_0 \chi(x'), \quad \eta_{CP} = \pm i .$$

# Standard Parametrization

$$U = V \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{pmatrix} \quad (4)$$

$$V = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13}e^{i\delta} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13}e^{i\delta} \end{pmatrix} \quad (5)$$

- $s_{ij} \equiv \sin \theta_{ij}$ ,  $c_{ij} \equiv \cos \theta_{ij}$ ,  $\theta_{ij} = [0, \frac{\pi}{2}]$ ,
- $\delta$  - Dirac CP-violation phase,  $\delta = [0, 2\pi]$ ,
- $\alpha_{21}$ ,  $\alpha_{31}$  - the two Majorana CP-violation phases.
- If  $\Delta m_{\odot}^2 = \Delta m_{21}^2 > 0$ ,  $\Delta m_{\text{atm}}^2 = \Delta m_{31}^2$ ,

then  $\theta_{12} = \theta_{\odot}$ ,  $\theta_{23} = \theta_{\text{atm}}$ ,  $\theta_{13} = \theta$ .

The angle  $\theta_{13}$  is limited by the data from the CHOOZ and Palo Verde experiments.

- $\alpha_{21}$ ,  $\alpha_{31}$ :

–  $\nu_l \leftrightarrow \nu_l$ ,  $\bar{\nu}_l \leftrightarrow \bar{\nu}_l$  **not sensitive**;

S.M. Bilenky, J. Hosek, S.T.P., 1980;  
P. Langacker, S.T.P., G. Steigman, S. Toshev, 1987

–  $|\langle m \rangle|$  in  $(\beta\beta)_{0\nu}$ -decay **depends** on  $\alpha_{21}$ ,  $\alpha_{31}$ ;

–  $\Gamma(\mu \rightarrow e + \gamma)$  etc. in SUSY theories **depend** on  $\alpha_{21,31}$ ;

– BAU, leptogenesis scenario:  $\alpha_{21,31}$  ?

# Neutrino Mixing Parameters

$$\theta_{12}, \theta_{23}, \theta_{13}$$

$\nu_j$

Dirac

Majorana

$$\delta$$

$$\delta, \alpha_{21}, \alpha_{31}$$

$$m_1, m_2, m_3$$

$m_1, m_2, m_3$  - in terms of  $\Delta m_{\odot}^2$ ,  $\Delta m_{\text{atm}}^2$  and  $\min(m_j)$

## Conventions

**A.**  $m_1 < m_2 < m_3$  (NH) or  $m_3 < m_1 < m_2$  (IH)

- $\Delta m_{\odot}^2 = \Delta m_{21}^2 > 0$

- $\Delta m_{\text{atm}}^2 = \Delta m_{31}^2 > 0$  (NH),  $\Delta m_{\text{atm}}^2 = \Delta m_{32}^2 < 0$  (IH)

- $m_2 = \sqrt{m_1^2 + \Delta m_{21}^2}$ ,  $m_3 = \sqrt{m_1^2 + \Delta m_{31}^2}$

**B.**  $m_1 < m_2 < m_3$

- $\Delta m_{\text{atm}}^2 = \Delta m_{31}^2 > 0$

Two possibilities:

$$\Delta m_{\odot}^2 = \Delta m_{21}^2 > 0, \text{ NH}$$

$$\Delta m_{\odot}^2 = \Delta m_{32}^2 > 0, \text{ IH}$$

“discrete” parameter

# Future Progress

- High precision determination of  $\Delta m_{\odot}^2$ ,  $\theta_{\odot}$ ,  $\Delta m_{\text{atm}}^2$ ,  $\theta_{\text{atm}}$ .
- Measurement of, or improving by at least a factor of (5 - 10) the existing upper limit on,  $\sin^2 \theta_{13}$ .
- Determination of the **type of the  $\nu$ - mass spectrum**

$$m_1 \ll m_2 \ll m_3, \quad \text{NH,}$$

$$m_3 \ll m_1 < m_2, \quad \text{IH,}$$

$$m_1 \cong m_2 \cong m_3, \quad m_{1,2,3}^2 \gg \Delta m_{\text{atm}}^2, \quad \text{QD; } m_j \gtrsim 0.20 \text{ eV.}$$

- Determining or obtaining significant constraints on the **absolute scale of neutrino masses**, or on  $\min(m_j)$ .
- Determination of the nature - **Dirac or Majorana**, of  $\nu_j$  .
- Status of the CP-symmetry in the lepton sector: **violated due to  $\delta$  (Dirac)**, and/or **due to  $\alpha_{21}$ ,  $\alpha_{31}$  (Majorana)**?
- Searching for possible manifestations, other than  $\nu_l$ -oscillations, of the non-conservation of  $L_l$ ,  $l = e, \mu, \tau$ , such as  **$\mu \rightarrow e + \gamma$ ,  $\tau \rightarrow \mu + \gamma$** , etc. decays.
- Understanding at fundamental level the mechanism giving rise to the  $\nu$ - masses and mixing and to the  $L_l$ -non-conservation, i.e., finding **The Theory of  $\nu$ -mixing**.

## $(\beta\beta)_{0\nu}$ –Decay Experiments:

- Majorana nature of  $\nu_j$
- Type of  $\nu$ –mass spectrum (NH, IH, QD)
- Absolute neutrino mass scale

${}^3\text{H}$   $\beta$ -decay, cosmology:  $m_\nu$  (QD, IH)

- CPV due to Majorana CPV phases

$\nu_j$ – Dirac or Majorana particles,  
**fundamental problem**

$\nu_j$ –Dirac: **conserved lepton charge exists,**

$$L = L_e + L_\mu + L_\tau, \quad \nu_j \neq \bar{\nu}_j$$

$\nu_j$ –Majorana: **no lepton charge is exactly conserved.**

$$\nu_j \equiv \bar{\nu}_j \quad - \quad \text{fermionic analogs of } \pi^0$$

The observed patterns of  $\nu$ –mixing and of  $\Delta m_{\text{atm}}^2$  and  $\Delta m_{\odot}^2$  can be related to Majorana  $\nu_j$  and an **approximate** symmetry:

$$L' = L_e - L_\mu - L_\tau$$

S.T.P., 1982

See-saw mechanism:  $\nu_j$ – Majorana

Establishing that  $\nu_j$  are Majorana particles would be as important as the discovery of  $\nu$ – oscillations.

If  $\nu_j$  – Majorana particles,  
 $U_{\text{PMNS}}$  contains (3- $\nu$  mixing)

$\delta$ -Dirac,  $\alpha_{21}, \alpha_{31}$  - Majorana physical CPV phases

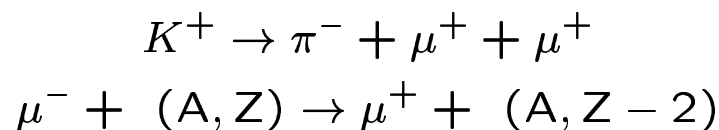
$\nu$ -oscillations  $\nu_l \leftrightarrow \nu_{l'}, \bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}, l, l' = e, \mu, \tau,$

- are not sensitive to the nature of  $\nu_j,$

S.M. Bilenky, J. Hosek, S.T.P., 1980;  
P. Langacker et al., 1987

- provide information on  $\Delta m_{jk}^2 = m_j^2 - m_k^2,$  but not on the absolute values of  $\nu_j$  masses.

The Majorana nature of  $\nu_j$  can manifest itself in the existence of  $\Delta L = \pm 2$  processes:



The process most sensitive to the possible Majorana nature of  $\nu_j$  -  $(\beta\beta)_{0\nu}$ -decay



of certain even-even nuclei,  $^{48}\text{Ca}, ^{76}\text{Ge}, ^{82}\text{Se}, ^{100}\text{Mo},$   
 $^{116}\text{Cd}, ^{130}\text{Te}, ^{136}\text{Xe}.$

$2n$  from  $(A, Z)$  exchange a virtual Majorana  $\nu_j$  (via the CC weak interaction) and transform into  $2p$  of  $(A, Z+2)$  and two free  $e^-$ .

$$A(\beta\beta)_{0\nu} \sim \langle m \rangle M(A,Z), \quad M(A,Z) - \text{NME},$$

$$|\langle m \rangle| = |m_1|U_{e1}|^2 + m_2|U_{e2}|^2 e^{i\alpha_{21}} + m_3|U_{e3}|^2 e^{i\alpha_{31}},$$

$\alpha_{21}, \alpha_{31}$  - the two Majorana CPVP of the PMNS matrix.

**CP-invariance:**  $\alpha_{21} = 0, \pm\pi, \alpha_{31} = 0, \pm\pi;$

$$\eta_{21} \equiv e^{i\alpha_{21}} = \pm 1, \quad \eta_{31} \equiv e^{i\alpha_{31}} = \pm 1$$

**relative CP-parities of  $\nu_1$  and  $\nu_2$ , and of  $\nu_1$  and  $\nu_3$  .**

L. Wolfenstein, 1981;

S.M. Bilenky, N. Nedelcheva, S.T.P., 1984;

B. Kayser, 1984.

Best sensitivity: Heidelberg-Moscow  $^{76}\text{Ge}$  experiment.

Claim for a positive signal at  $> 3\sigma$ :

H. Klapdor-Kleingrothaus et al., PL B586 (2004),

$$|\langle m \rangle| = (0.1 - 0.9) \text{ eV (99.73\% C.L.)}.$$

IGEX  $^{76}\text{Ge}$ :  $|\langle m \rangle| < (0.33 - 1.35) \text{ eV (90\% C.L.)}.$

Taking data - NEMO3 ( $^{100}\text{Mo}$ ), CUORICINO ( $^{130}\text{Te}$ ):

$$|\langle m \rangle| < (0.7-1.2) \text{ eV, } |\langle m \rangle| < (0.2-1.1) \text{ eV (90\% C.L.)}.$$

**Large number of projects:**  $|\langle m \rangle| \sim (0.01 - 0.05) \text{ eV}$

CUORE -  $^{130}\text{Te}$ ,

GERDA -  $^{76}\text{Ge}$ ,

EXO -  $^{136}\text{Xe}$ ,

MAJORANA -  $^{76}\text{Ge}$ ,

MOON -  $^{100}\text{Mo}$ ,

CANDLES -  $^{48}\text{Ca}$ ,

XMASS -  $^{136}\text{Xe}$ .



$$|\langle m \rangle| : m_j, |U_{ej}|^2, \alpha_{21,31}$$

$m_{1,2,3}$  - in terms of  $\min(m_j)$ ,  $\Delta m_{\text{atm}}^2$ ,  $\Delta m_{\odot}^2$

$|U_{ej}|^2$  - in terms of  $\sin^2 \theta_{\odot}$ ,  $\cos^2 \theta_{\odot}$ ,  $\sin^2 \theta$  (CHOOZ)

In convention (B) in which always  $m_1 < m_2 < m_3$ ,

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2, \quad m_3 = \sqrt{m_1^2 + \Delta m_{\text{atm}}^2},$$

while either

$$\Delta m_{\odot}^2 \equiv \Delta m_{21}^2 \quad (\text{normal mass hierarchy}) \quad \text{or}$$

$$\Delta m_{\odot}^2 \equiv \Delta m_{32}^2 \quad (\text{inverted mass hierarchy}).$$

Normal mass hierarchy:  $m_2 = \sqrt{m_1^2 + \Delta m_{\odot}^2}$ ,

$$|U_{e1}|^2 = \cos^2 \theta_{\odot} (1 - |U_{e3}|^2), \quad |U_{e2}|^2 = \sin^2 \theta_{\odot} (1 - |U_{e3}|^2),$$

$$|U_{e3}|^2 \equiv \sin^2 \theta \quad (\text{CHOOZ})$$

Inverted mass hierarchy:  $m_2 = \sqrt{m_1^2 + \Delta m_{\text{atm}}^2 - \Delta m_{\odot}^2}$ ,

$$|U_{e2}|^2 = \cos^2 \theta_{\odot} (1 - |U_{e1}|^2), \quad |U_{e3}|^2 = \sin^2 \theta_{\odot} (1 - |U_{e1}|^2),$$

$$|U_{e1}|^2 \equiv \sin^2 \theta \quad (\text{CHOOZ}).$$

The neutrino mass spectrum –

*normal hierarchical (NH)* if  $m_1 \ll m_2 \ll m_3$ ,

*inverted hierarchical (IH)* if  $m_1 \ll m_2 \cong m_3$ .

The *quasi-degenerate (QD)* mass spectrum is realized if  $m_1 \cong m_2 \cong m_3$ ,  $m_{1,2,3}^2 \gg \Delta m_{\text{atm}}^2$ .

Given  $\Delta m_{\text{atm}}^2$ ,  $\Delta m_{\odot}^2$ ,  $\theta_{\odot}$ ,  $\theta$ ,

$$|\langle m \rangle| = |\langle m \rangle| (m_{\text{min}}, \alpha_{21}, \alpha_{31}; S), \quad S = NH, IH.$$

## Oscillation Parameters

$$\Delta m_{\odot}^2 = 8.0 \times 10^{-5} \text{ eV}^2, \quad 3\sigma(\Delta m_{\odot}^2) = 12\%,$$

$$\sin^2 \theta_{\odot} = 0.31, \quad 3\sigma(\sin^2 \theta_{\odot}) = 24\%,$$

$$|\Delta m_{\text{atm}}^2| = 2.2 \times 10^{-3} \text{ eV}^2, \quad 3\sigma(|\Delta m_{\text{atm}}^2|) = 50\%.$$

### Future:

SNO III:  $3\sigma(\sin^2 \theta_{\odot}) = 21\%$  ;

3 kTy KamLAND:  $3\sigma(\Delta m_{\odot}^2) = 7\%$  ,  $3\sigma(\sin^2 \theta_{\odot}) = 18\%$  ;  
A. Bandyopadhyay et al., hep-ph/0410283

SK-Gd (0.1% Gd:  $43 \times (\text{KL } \bar{\nu}_e \text{ rate})$ ), 3y:  $3\sigma(\Delta m_{\odot}^2) \cong 4\%$   
S. Choubey, S.T.P., hep-ph/0404103;  
J. Beacom and M. Vagins, hep-ph/0309300

KL type reactor  $\bar{\nu}_e$  detector,  $L \sim 60 \text{ km}$ ,  $\sim 60 \text{ GW kTy}$ :

$$3\sigma(\sin^2 \theta_{\odot}) \cong 12\%$$

A. Bandyopadhyay et al., hep-ph/0410283 and hep-ph/0302243;  
H. Minakata et al., hep-ph/0407326

T2K (SK):  $3\sigma(|\Delta m_{\text{atm}}^2|) \cong 12\%$

P. Huber et al., hep-ph/0403068

$\text{sgn}(\Delta m_{\text{atm}}^2)$ : atmospheric  $\nu$  experiments, studying the sub-dominant  $\nu_{\mu(e)} \rightarrow \nu_{e(\mu)}$  and  $\bar{\nu}_{\mu(e)} \rightarrow \bar{\nu}_{e(\mu)}$  oscillations; LBL  $\nu$ -oscillation experiments (T2K, NO $\nu$ A);  $\nu$ -factory.

$\sin^2 \theta_{13}$ : reactor  $\bar{\nu}_e$  experiments,  $L \sim (1 - 2) \text{ km}$ : Double CHOOZ, Braidwood, Daya-Bay, KASKA - factor (5 - 10).

# Absolute Neutrino Mass Measurements

The Troitzk and Mainz  ${}^3\text{H}$   $\beta$ -decay experiments

$$m_{\nu_e} < 2.2 \text{ eV} \quad (95\% \text{ C.L.})$$

There are prospects to reach sensitivity

$$\text{KATRIN :} \quad m_{\nu_e} \sim 0.2 \text{ eV}$$

Cosmological and astrophysical data: the WMAP result combined with data from large scale structure surveys (2dFGRS, SDSS)

$$\sum_j m_j \equiv \Sigma < 0.70 \text{ eV} \quad (95\% \text{ C.L.}) \quad (X \sim 3)$$

The WMAP and future PLANCK experiments can be sensitive to

$$\sum_j m_j \cong 0.4 \text{ eV}$$

Data on weak lensing of galaxies by large scale structure, combined with data from the WMAP and PLANCK experiments may allow to determine

$$\sum_j m_j : \quad \delta \cong 0.04 \text{ eV.}$$

**Normal Hierarchical Spectrum,  $m_1 \ll m_2 \ll m_3$ :**

$$m_2 \cong \sqrt{\Delta m_{\odot}^2} \cong (8.4 - 9.4) \times 10^{-3} \text{ eV} \quad (3\sigma),$$

$$m_3 \cong \sqrt{\Delta m_{\text{atm}}^2} \cong (3.7 - 5.7) \times 10^{-2} \text{ eV} \quad (3\sigma).$$

**Inverted Hierarchical Spectrum,  $m_3 \ll m_1 \cong m_2$ :**

$$m_{1,2} \cong \sqrt{|\Delta m_{\text{atm}}^2|} \cong (3.7 - 5.7) \times 10^{-2} \text{ eV} .$$

**Quasi-Degenerate Spectrum,  $m_1 \cong m_2 \cong m_3 \equiv m$ :**

$$m_{1,2,3}^2 \gg |\Delta m_{\text{atm}}^2| .$$

**Using  $|\Delta m_{\text{atm}}^2|$  and  $\Delta m_{\odot}^2$  inferred from the data one has**

**Normal (Inverted) Hierarchical Spectrum for**

$$m_1 \ll 0.02 \text{ eV} \quad (m_3 < 0.02 \text{ eV}) ;$$

**Spectrum with Partial Hierarchy for**

$$0.02 \text{ eV} \lesssim m_{1(3)} \lesssim 0.20 \text{ eV} ;$$

**Quasi-Degenerate Spectrum for**

$$m_{1(3)} \gtrsim 0.20 \text{ eV} .$$

# Hierarchical Neutrino Mass Spectrum

$$m_1 \ll m_2 \ll m_3.$$

The pattern corresponds to

$$m_1 \ll \sqrt{\Delta m_{21}^2} \ll \sqrt{\Delta m_{32}^2}.$$

It is possible to identify

$$\Delta m_{\odot}^2 \equiv \Delta m_{21}^2, \quad \Delta m_{\text{atm}}^2 \equiv \Delta m_{32}^2,$$

$$|U_{e1}|^2 = \cos^2 \theta_{\odot} (1 - |U_{e3}|^2), \quad |U_{e2}|^2 = \sin^2 \theta_{\odot} (1 - |U_{e3}|^2),$$

$$|U_{e3}|^2 \equiv \sin^2 \theta_{13} < 0.05 \quad (\text{CHOOZ} + \nu_A + \nu_{\odot} + \text{KL}).$$

This implies:

$$m_2 \simeq \sqrt{\Delta m_{\odot}^2}, \quad m_3 \simeq \sqrt{\Delta m_{\text{atm}}^2}.$$

One has

$$\begin{aligned} |\langle m \rangle| &= \left| (m_1 \cos^2 \theta_{\odot} + \sqrt{\Delta m_{\odot}^2} \sin^2 \theta_{\odot}) (1 - |U_{e3}|^2) e^{i\alpha_{21}} \right. \\ &\quad \left. + \sqrt{\Delta m_{\text{atm}}^2} |U_{e3}|^2 e^{i\alpha_{31}} \right| \\ &\simeq \left| \sqrt{\Delta m_{\odot}^2} (1 - |U_{e3}|^2) \sin^2 \theta_{\odot} + \sqrt{\Delta m_{\text{atm}}^2} |U_{e3}|^2 e^{i(\alpha_{31} - \alpha_{21})} \right| \end{aligned}$$

Even if  $m_1 = 0$ ,  $|\langle m \rangle|$  depends on Majorana CPV phase  $\alpha_{32} = \alpha_{31} - \alpha_{21}$ . Reflects the fact that in contrast to the case of massive Dirac  $\nu$ 's (or quarks), CPV can take place in the mixing of only 2 massive Majorana  $\nu$ 's.

# Inverted Hierarchical $\nu$ -Mass Spectrum

$$m_3 \ll m_1 \simeq m_2.$$

Can also be defined by

$$m_3 \ll (<) \sqrt{\Delta m_{21}^2} \ll \sqrt{\Delta m_{23}^2}.$$

We can identify

$$\Delta m_{\odot}^2 \equiv \Delta m_{21}^2, \quad \Delta m_{\text{atm}}^2 \equiv \Delta m_{32}^2 \simeq \Delta m_{31}^2,$$

$$|U_{e3}|^2 = \sin^2 \theta_{13} < 0.05 \quad (\text{CHOOZ} + \nu_A + \nu_{\odot} + \text{KL}),$$

$$|U_{e1}|^2 = \cos^2 \theta_{\odot} (1 - |U_{e3}|^2), \quad |U_{e2}|^2 = \sin^2 \theta_{\odot} (1 - |U_{e3}|^2),$$

$$m_1 \simeq m_2 \simeq \sqrt{|\Delta m_{\text{atm}}^2|}.$$

Neglecting  $m_3 |U_{e3}^2|$  in  $|\langle m \rangle|$ ,

$$|\langle m \rangle| \simeq \sqrt{|\Delta m_{\text{atm}}^2| (1 - |U_{e3}|^2)} \sqrt{1 - \sin^2 2\theta_{\odot} \sin^2 \left( \frac{\alpha_{21}}{2} \right)},$$

$$\sqrt{|\Delta m_{\text{atm}}^2| (1 - |U_{e3}|^2)} |\cos 2\theta_{\odot}| \leq |\langle m \rangle| \leq \sqrt{|\Delta m_{\text{atm}}^2| (1 - |U_{e3}|^2)}.$$

The upper and lower limits - CP-conserving:

$$\alpha_{21} = 0, \quad \alpha_{21} = \pm\pi.$$

$$\sin^2 \frac{\alpha_{21}}{2} = \left( 1 - \frac{|\langle m \rangle|^2}{|\Delta m_{\text{atm}}^2| (1 - |U_{e3}|^2)^2} \right) \frac{1}{\sin^2 2\theta_{\odot}}.$$

# Three Quasi-Degenerate Neutrinos

$$m_1 \simeq m_2 \simeq m_3 \equiv m, \quad m^2 \gg |\Delta m_{\text{atm}}^2|.$$

As for the NH neutrino mass spectrum we have:

$$\Delta m_{\odot}^2 \equiv \Delta m_{21}^2, \quad \Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2,$$

$$|U_{e1}|^2 = \cos^2 \theta_{\odot} (1 - |U_{e3}|^2), \quad |U_{e2}|^2 = \sin^2 \theta_{\odot} (1 - |U_{e3}|^2),$$

$$|U_{e3}|^2 = \sin^2 \theta_{13} < 0.05 \quad (\text{CHOOZ} + \nu_A + \nu_{\odot} + \text{KL}).$$

The mass scale  $m$  effectively coincides with the  $\bar{\nu}_e$  mass  $m_{\bar{\nu}_e}$  measured in the current  ${}^3\text{H}$   $\beta$ -decay experiments:

$$m \cong m_{\bar{\nu}_e}.$$

Thus,  $m < 2.2$  eV. Cosmology:  $m \lesssim 0.6$  eV.

The QD spectrum - realized for  $m$ , which can be measured in the  ${}^3\text{H}$   $\beta$ -decay experiment KATRIN,  $m_{\bar{\nu}_e} \gtrsim (0.2 - 0.3)$  eV.

The effective Majorana mass  $|\langle m \rangle|$  is given by

$$\begin{aligned} |\langle m \rangle| &\cong m \left| \cos^2 \theta_{\odot} (1 - |U_{e3}|^2) \right. \\ &\quad \left. + \sin^2 \theta_{\odot} (1 - |U_{e3}|^2) e^{i\alpha_{21}} + |U_{e3}|^2 e^{i\alpha_{31}} \right| \\ &\cong m \left| \cos^2 \theta_{\odot} + \sin^2 \theta_{\odot} e^{i\alpha_{21}} \right|; \end{aligned}$$

$$m |\cos 2\theta_{\odot}| \lesssim |\langle m \rangle| \lesssim m; \quad \text{limits: } \alpha_{21} = 0; \pm \pi - \text{CPC}$$

$$\sin^2 \frac{\alpha_{21}}{2} \cong \left( 1 - \frac{|\langle m \rangle|^2}{m(1 - |U_{e3}|^2)^2} \right) \frac{1}{\sin^2 2\theta_{\odot}}.$$

## Solar neutrino and KamLAND data:

$\cos 2\theta_{\odot} = 0.0$  excluded at  $> 6$  s.d.

Best fit value:  $\cos 2\theta_{\odot} \simeq 0.40$

$\cos 2\theta_{\odot} \gtrsim 0.28$ , 95% C.L.

Normal hierarchical spectrum:

$$(|\langle m \rangle|)_{\max} \lesssim 0.005 \text{ eV}$$

Inverted hierarchical spectrum:

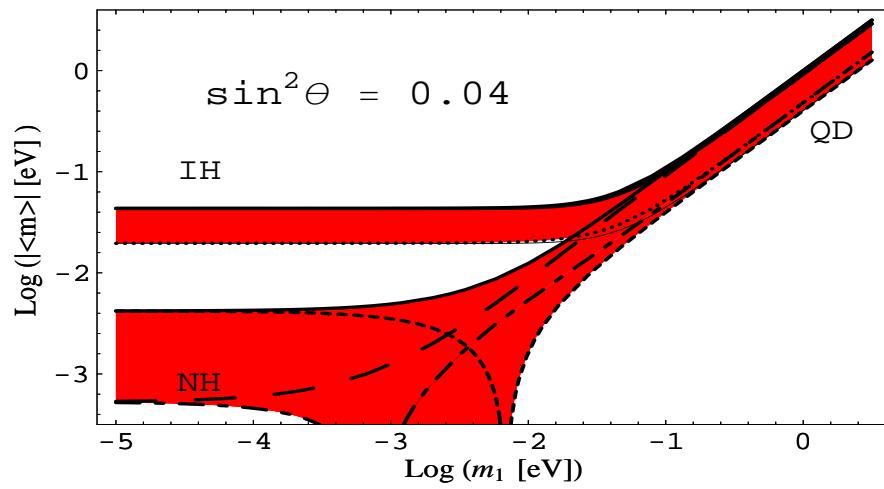
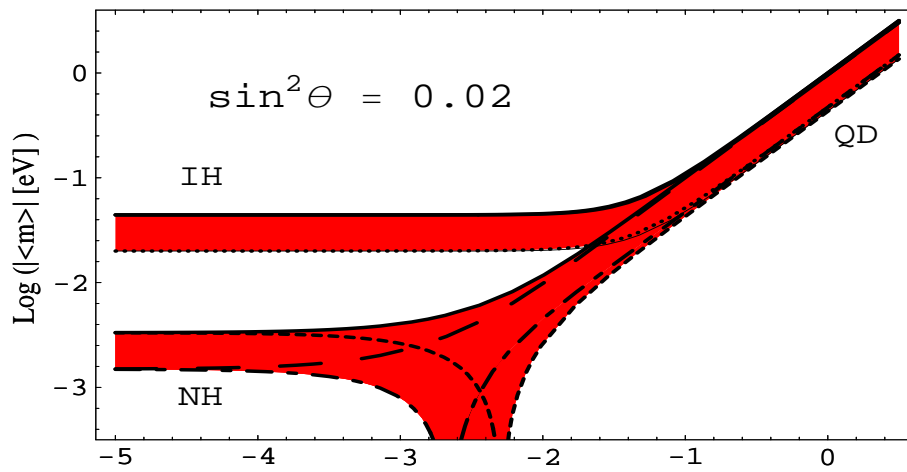
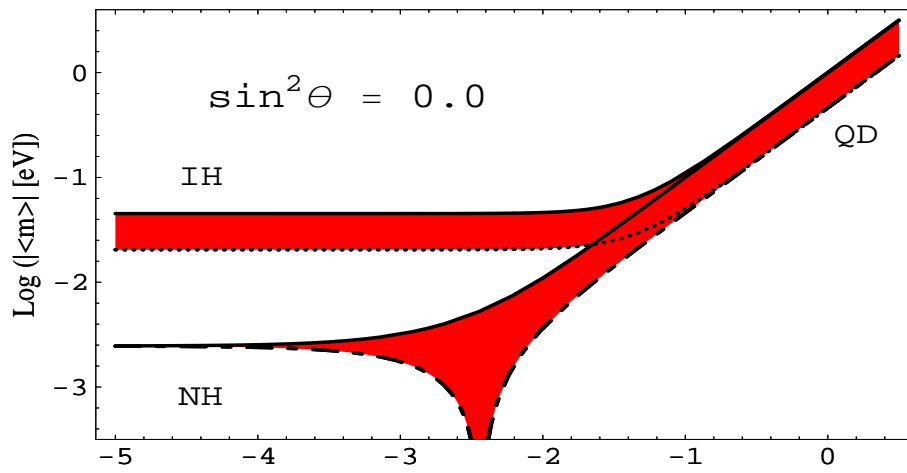
$$(|\langle m \rangle|)_{\min} \simeq \sqrt{|\Delta m_{\text{atm}}^2|} \cos 2\theta_{\odot} \cos^2 \theta_{13} \gtrsim 0.01 \text{ eV}$$

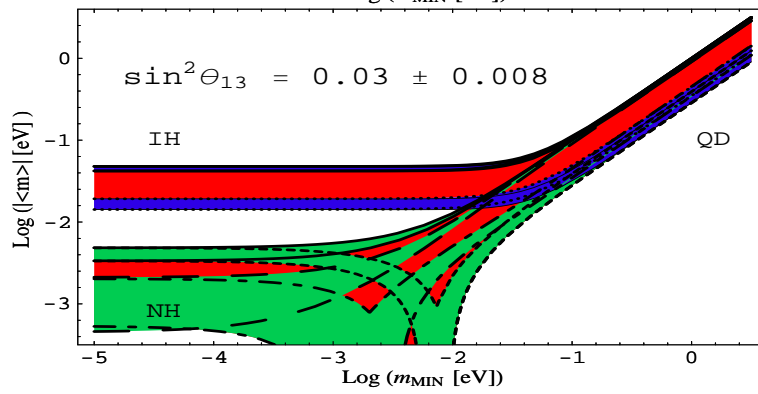
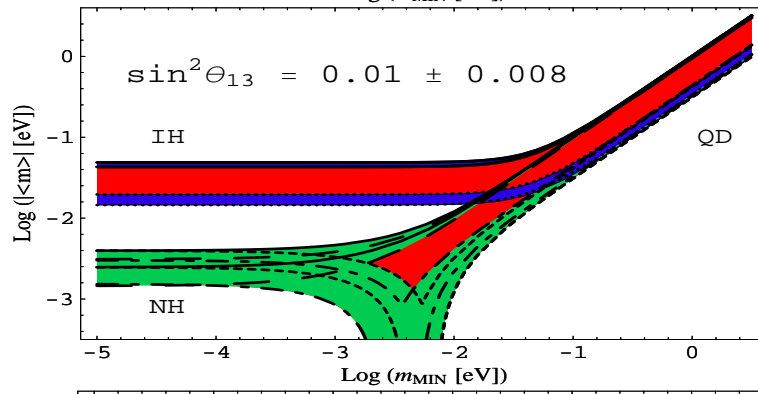
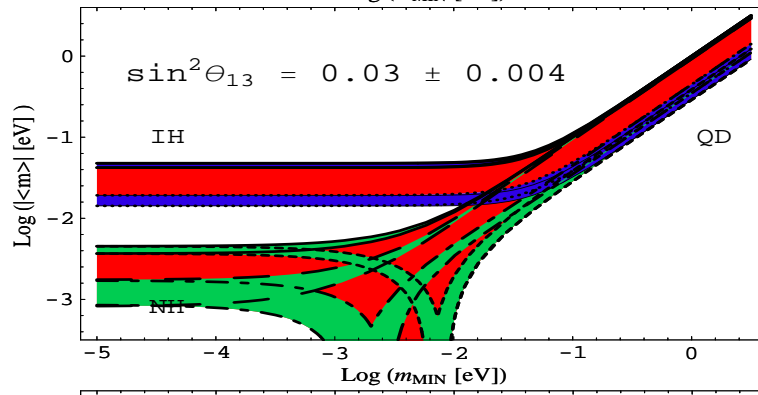
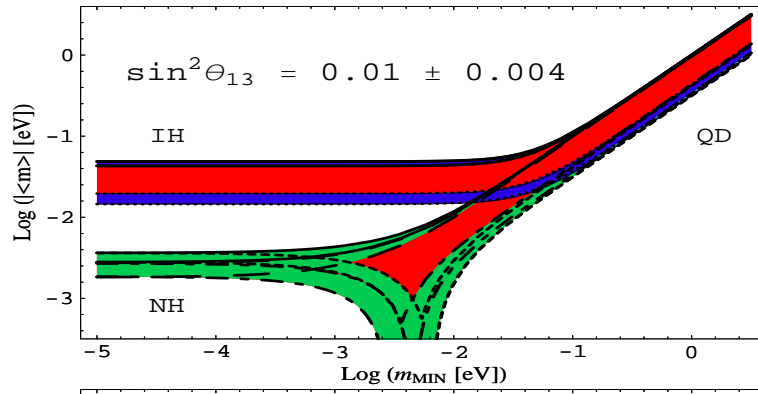
$$(|\langle m \rangle|)_{\max} \simeq \sqrt{|\Delta m_{\text{atm}}^2|} \cos^2 \theta_{13} \lesssim 0.055 \text{ eV}$$

Quasi-degenerate spectrum:

$$(|\langle m \rangle|)_{\min} \simeq m (\cos 2\theta_{\odot} \cos^2 \theta_{13} - \sin^2 \theta_{13}) \gtrsim 0.06 \text{ eV}$$







S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226

# Nuclear Matrix Element Uncertainty

$$|\langle m \rangle| = \zeta ( (|\langle m \rangle|_{\text{exp}})_{\text{MIN}} \pm \Delta) , \quad \zeta \geq 1,$$

$(|\langle m \rangle|_{\text{exp}})_{\text{MIN}}$  – obtained with the **maximal physically allowed value of NME**.

A measurement of the  $(\beta\beta)_{0\nu}$ -decay half-life time

$$(|\langle m \rangle|_{\text{exp}})_{\text{MIN}} - \Delta \leq |\langle m \rangle| \leq \zeta ( (|\langle m \rangle|_{\text{exp}})_{\text{MIN}} + \Delta) .$$

The currently estimated range of  $\zeta^2$ :

$${}^{48}\text{Ca}, \quad \zeta^2 \simeq 3.5$$

$${}^{76}\text{Ge}, \quad \zeta^2 \simeq 10$$

$${}^{82}\text{Se}, \quad \zeta^2 \simeq 10$$

$${}^{130}\text{Te}, \quad \zeta^2 \simeq 38.7$$

S. Elliot, P. Vogel, 2002

NH vs IH (QD):

$$\zeta |\langle m \rangle|_{\text{max}}^{\text{NH}} < |\langle m \rangle|_{\text{min}}^{\text{IH(QD)}} , \quad \zeta \geq 1 .$$

IH vs QD:

$$\zeta |\langle m \rangle|_{\text{max}}^{\text{IH}} < |\langle m \rangle|_{\text{min}}^{\text{QD}} , \quad \zeta \geq 1 .$$

These conditions imply upper limits on  $\tan^2 \theta_{\odot}$ , which are functions of the neutrino oscillation parameters and of  $\zeta$ .

S. Pascoli, S.T.P., W. Rodejohann, 2003

## Method of Analysis

$$\Gamma_{\text{th}} = G |\mathcal{M}|^2 (|\langle m \rangle|(\mathbf{x}))^2, \quad \mathbf{x} = (\mathbf{x}_{\text{osc}}, \mathbf{x}_{\beta\beta}^{0\nu})$$

$$\mathbf{x}_{\text{osc}} = (\theta_{12}, \theta_{13}, |\Delta \mathbf{m}_{31}^2|, \Delta \mathbf{m}_{21}^2),$$

$$\mathbf{x}_{\beta\beta}^{0\nu} = (m_0, \text{sgn}(\Delta \mathbf{m}_{31}^2), \alpha_{21}, \alpha_{31}).$$

$$|\langle m \rangle|^{\text{obs}} \equiv \sqrt{\frac{\Gamma_{\text{obs}}}{G}} \frac{1}{|\mathcal{M}_0|}, \quad \sigma_{\beta\beta} = \frac{1}{2} \frac{1}{\sqrt{\Gamma_{\text{obs}} G}} \frac{1}{|\mathcal{M}_0|} \sigma(\Gamma_{\text{obs}}),$$

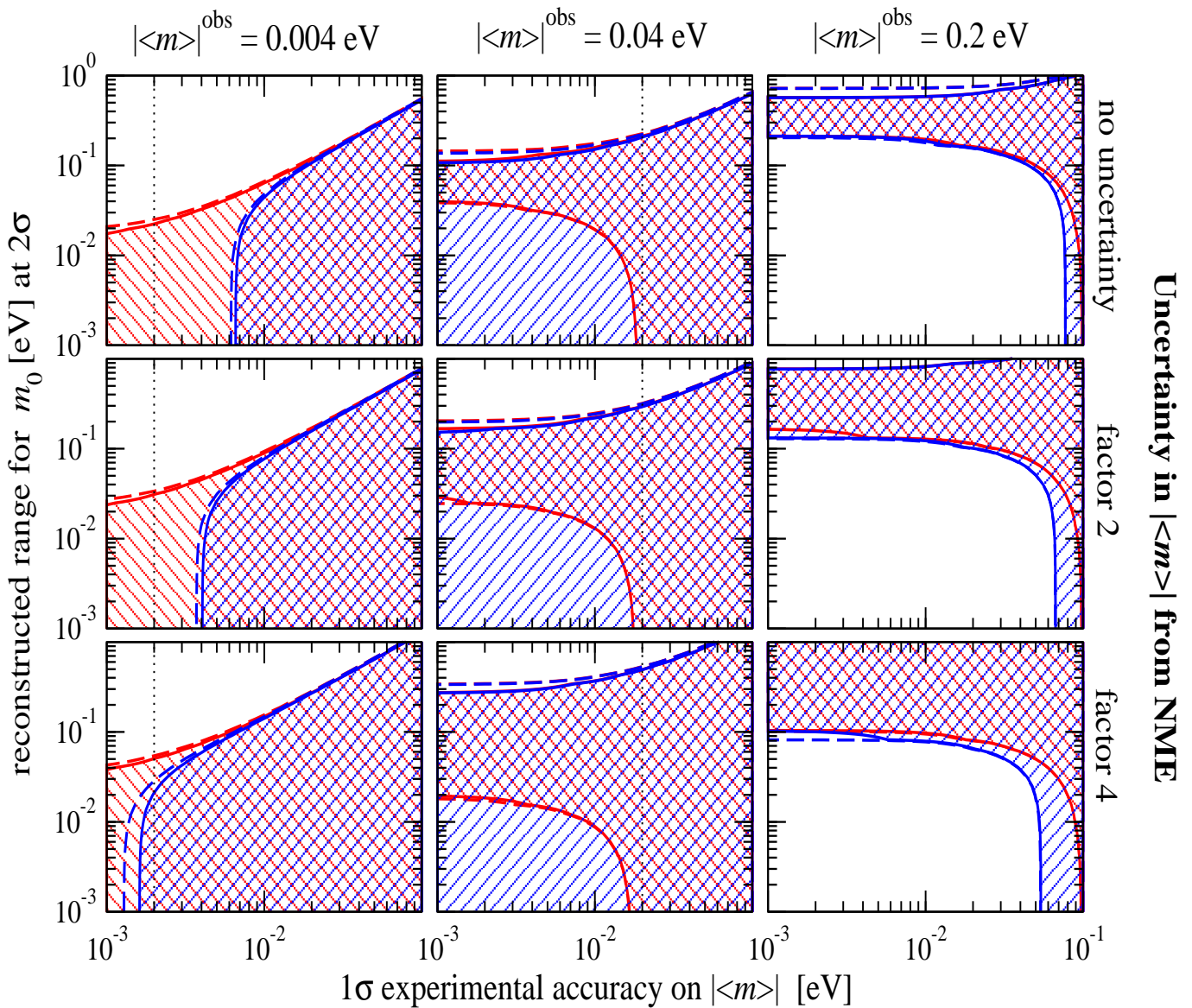
$|\mathcal{M}_0|$  is some nominal value of the NME.

$$\chi^2(\mathbf{x}_{\beta\beta}^{0\nu}, \mathbf{F}) = \min_{\xi \in [1/\sqrt{F}, \sqrt{F}]} \frac{[\xi |\langle m \rangle|(\mathbf{x}) - |\langle m \rangle|^{\text{obs}}]^2}{\sigma_{\beta\beta}^2 + \xi^2 \sigma_{\text{th}}^2}.$$

$$\xi \equiv \frac{|\mathcal{M}|}{|\mathcal{M}_0|}, \quad \xi = [1/\sqrt{F}, \sqrt{F}], \quad F \geq 1,$$

$|\mathcal{M}|$  is the *true* value of the NME.

# Absolute Neutrino Mass Scale



normal ordering

inverted ordering

**dashed:**  $\sigma(\sin^2\theta_{13}) = 0.016$ ,  $\sigma(\sin^2\theta_{12}) = 7.5\%$ ,  $\sigma(\Delta m_{21}^2) = 4\%$ ,  $\sigma(\Delta m_{31}^2) = 13\%$

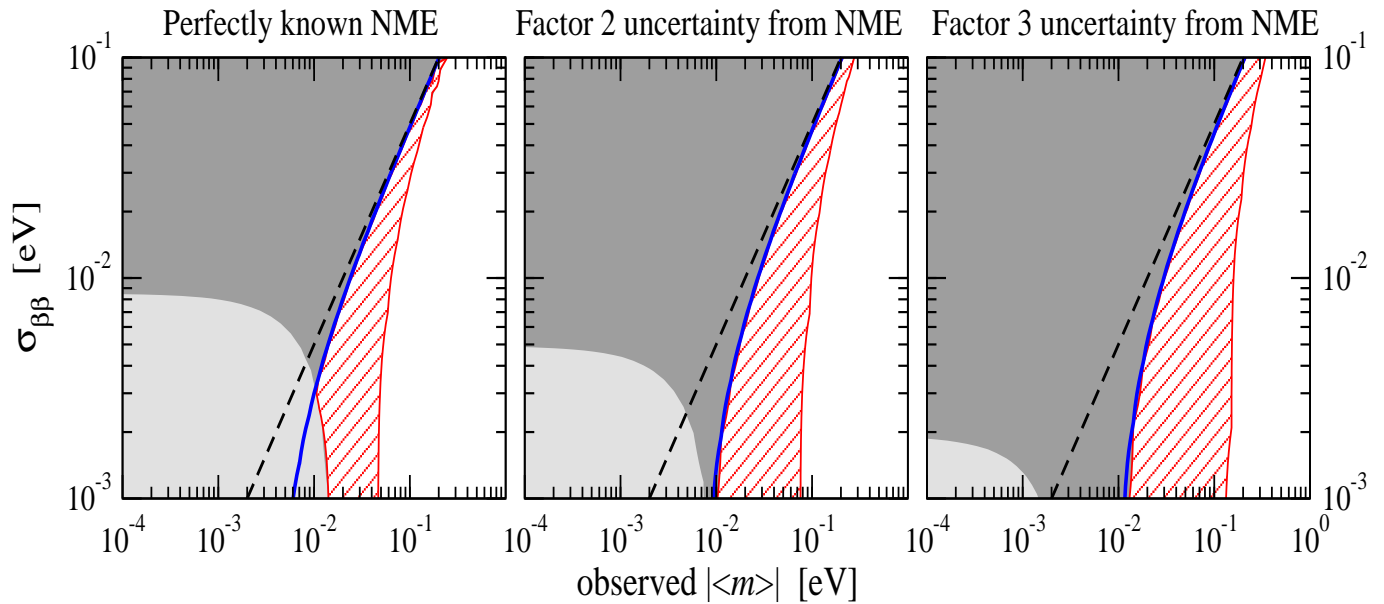
$\sin^2\theta_{13} = 0$

**solid:**  $\sigma(\sin^2\theta_{13}) = 0.002$ ,  $\sigma(\sin^2\theta_{12}) = 3.0\%$ ,  $\sigma(\Delta m_{21}^2) = 2\%$ ,  $\sigma(\Delta m_{31}^2) = 5\%$

$\sin^2\theta_{12} = 0.31$

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226

# Distinguishing Between Different Spectra



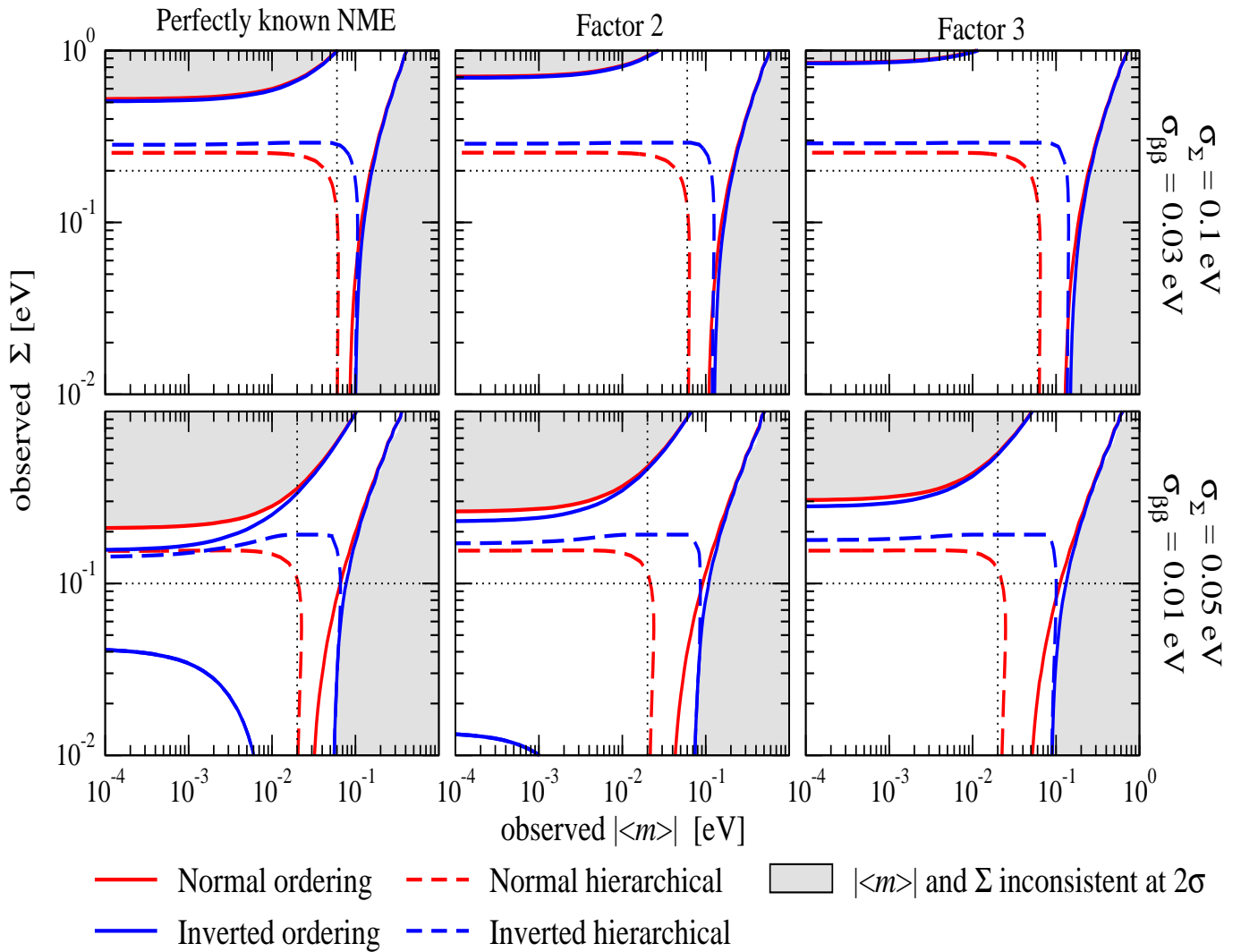
- No information on the mass ordering
- Inverted ordering excluded at  $2\sigma$
- To the right a signal is observed at  $2\sigma$
- Either IH or QD spectrum
- QD with no information on ordering
- To the right the NH spectrum is excluded

$$\sin^2 \theta_{13} = 0.03 \pm 0.006, \quad \sin^2 \theta_{12} = 0.31 \pm 3\%, \quad \Delta m_{21}^2 = 8 \times 10^{-5} \pm 2\%, \quad |\Delta m_{31}^2| = 2.2 \times 10^{-3} \pm 3\%$$

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226

# Distinguishing Between Different Spectra

Uncertainty in  $\langle m \rangle$  from NME:



$$\sin^2 \theta_{13} = 0 \pm 0.002, \quad \sin^2 \theta_{12} = 0.31 \pm 3\%, \quad \Delta m_{21}^2 = 8 \times 10^{-5} \pm 2\%, \quad |\Delta m_{31}^2| = 2.2 \times 10^{-3} \pm 3\%$$

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226

# Majorana CPV Phases and $|\langle m \rangle|$

IH spectrum:  $m_{\min} < 0.01$  eV,  $\sin^2 \theta_{\odot}$  – negligible

$$\sqrt{|\Delta m_{\text{atm}}^2|} |\cos 2\theta_{\odot}| \leq |\langle m \rangle| \leq \sqrt{|\Delta m_{\text{atm}}^2|}$$

“Just CP-violating” region:

$$(|\langle m \rangle|_{\text{exp}})_{\text{MAX}} < \sqrt{(\Delta m_{\text{atm}}^2)_{\text{MIN}}},$$

$$(|\langle m \rangle|_{\text{exp}})_{\text{MIN}} > \sqrt{(\Delta m_{\text{atm}}^2)_{\text{MAX}} (\cos 2\theta_{\odot})_{\text{MAX}}},$$

$$|\langle m \rangle| = \zeta (|\langle m \rangle|_{\text{exp}})_{\text{MIN}} \pm \Delta, \quad \zeta \geq 1$$

Necessary condition for establishing CP-violation:

$$1 \leq \zeta < \frac{\sqrt{(\Delta m_{\text{atm}}^2)_{\text{MIN}}}}{\sqrt{(\Delta m_{\text{atm}}^2)_{\text{MAX}} (\cos 2\theta_{\odot})_{\text{MAX}} + 2\Delta}} \simeq \frac{1}{(\cos 2\theta_{\odot})_{\text{MAX}}}$$

QD spectrum,  $m_{1,2,3} \simeq m_0 \gtrsim 0.20$  eV,  $\sin^2 \theta_{13}$  negligible - similar condition:  $\Delta m_{\text{atm}}^2 \rightarrow m_0^2$ .

CPV can be established provided

- $|\langle m \rangle|$  measured with  $\Delta \lesssim 15\%$  ;
- $\Delta m_{\text{atm}}^2$  (IH) or  $m_0$  (QD) measured with  $\delta \lesssim 10\%$  ;
- $\xi < 2$  ( $\xi \lesssim 1.5$ ) ;
- $\alpha_{21}$  (QD): in the interval  $\sim [\frac{\pi}{4} - \frac{3\pi}{4}]$ , or  $\sim [\frac{5\pi}{4} - \frac{3\pi}{2}]$  ;
- $\tan^2 \theta_{\odot} \gtrsim 0.44$  .

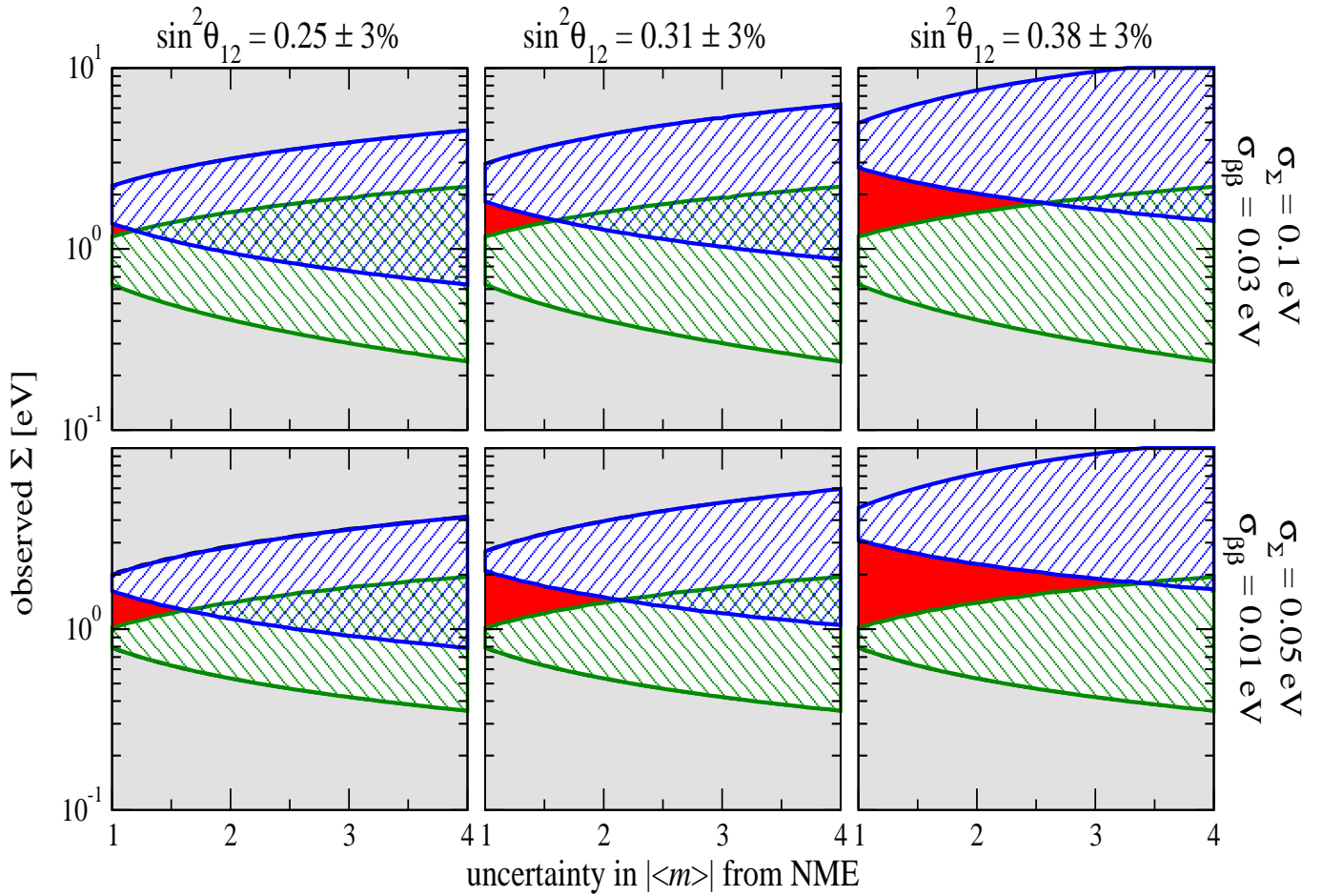
S. Pascoli, S.T.P., W. Rodejohann, 2002

S. Pascoli, S.T.P., L. Wolfenstein, 2002


“No-go for detecting CP-Violation via  $(\beta\beta)_{0\nu}$ -decay”


V. Barger *et al.*, 2002






 data consistent with  $\alpha_{21} = \pi$

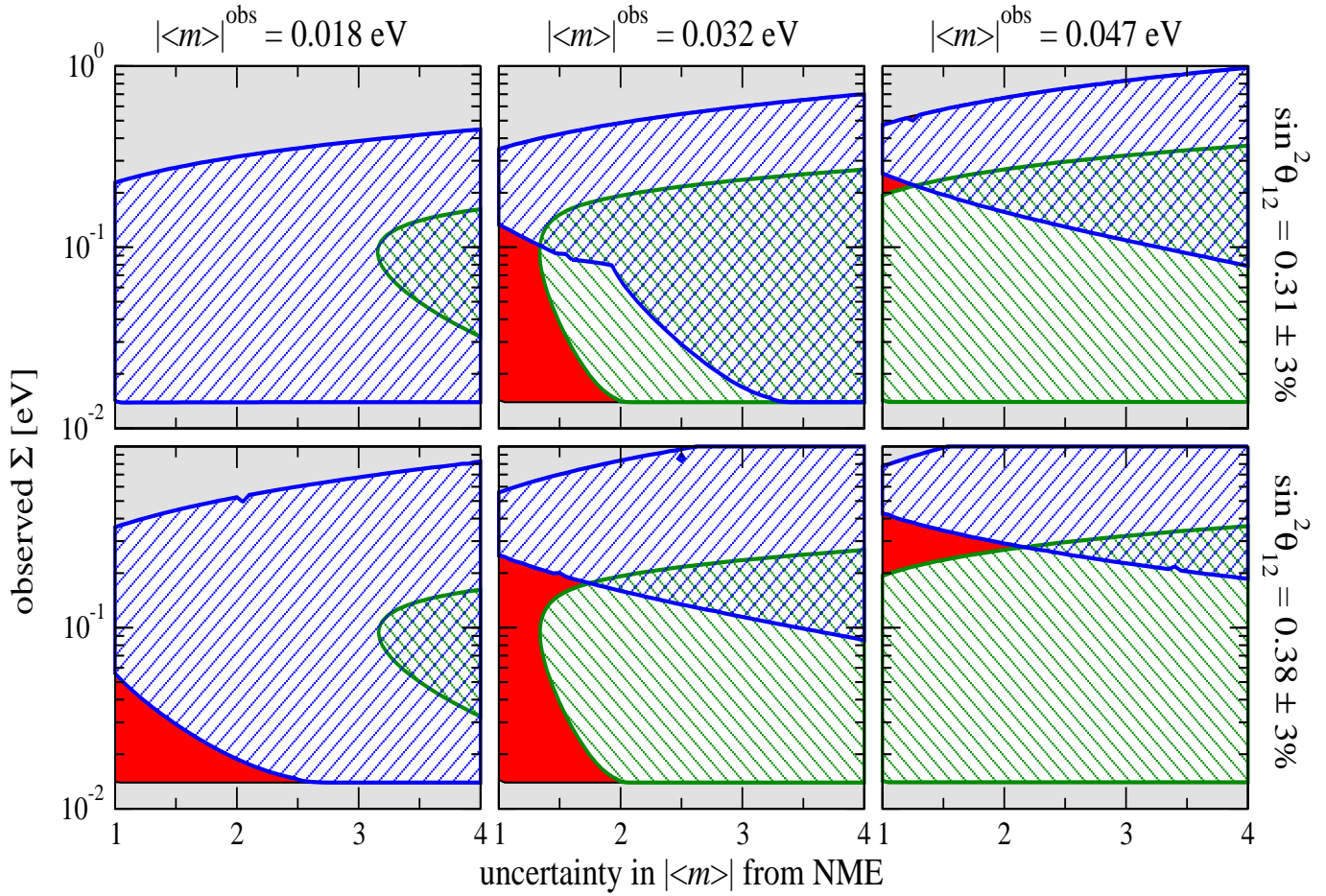
 data consistent with  $\alpha_{21} = 0$

  $|\langle m \rangle|$  and  $\Sigma$  inconsistent at  $2\sigma$

 CP violation established at  $2\sigma$

$\sin^2\theta_{13} = 0 \pm 0.002$ ,  $\Delta m_{21}^2 = 8 \times 10^{-5} \pm 2\%$ ,  $\Delta m_{31}^2 = 2.2 \times 10^{-3} \pm 3\%$       observed  $|\langle m \rangle| = 0.3$  eV

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226



data consistent with  $\alpha_{21} = \pi$

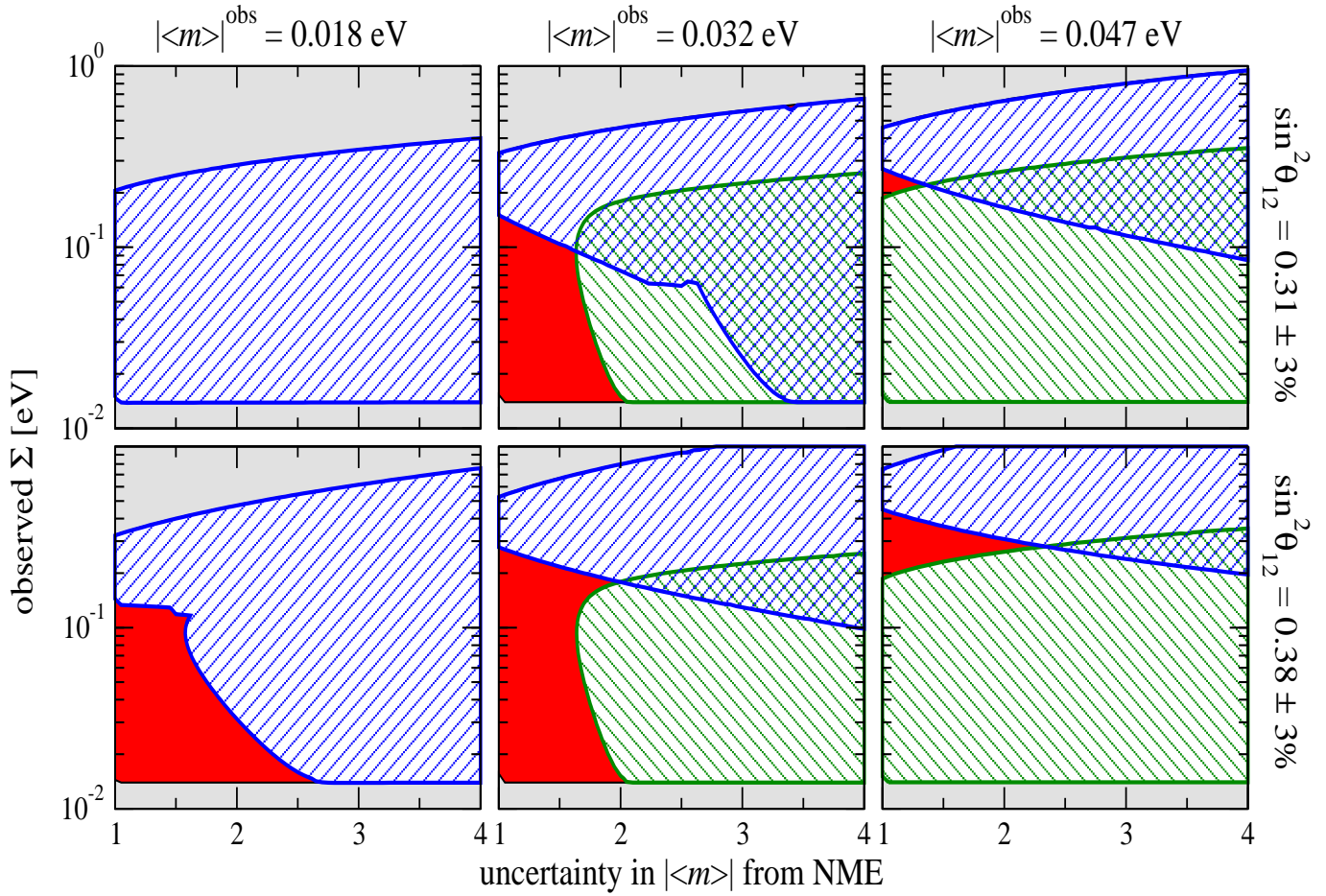
data consistent with  $\alpha_{21} = 0$

$|\langle m \rangle|$  and  $\Sigma$  inconsistent at  $2\sigma$

CP violation established at  $2\sigma$

$$\sin^2 \theta_{13} = 0 \pm 0.002, \quad \Delta m_{21}^2 = 8 \times 10^{-5} \pm 2\%, \quad \Delta m_{31}^2 = -2.2 \times 10^{-3} \pm 3\%, \quad \sigma_{\beta\beta} = 0.004 \text{ eV}, \quad \sigma_{\Sigma} = 0.04 \text{ eV}$$

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226



data consistent with  $\alpha_{21} = \pi$

data consistent with  $\alpha_{21} = 0$

$|\langle m \rangle|$  and  $\Sigma$  inconsistent at  $2\sigma$

CP violation established at  $2\sigma$

$$\sin^2 \theta_{13} = 0 \pm 0.002, \quad \Delta m_{21}^2 = 8 \times 10^{-5} \pm 2\%, \quad \Delta m_{31}^2 = -2.2 \times 10^{-3} \pm 3\%, \quad \sigma_{\beta\beta} = 0.002 \text{ eV}, \quad \sigma_{\Sigma} = 0.04 \text{ eV}$$

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226

# On the NME Uncertainties

The  $(\beta\beta)_{0\nu}$ -decay half-life

$$(T_{1/2}^{0\nu}(A, Z))^{-1} = |\langle m \rangle|^2 |M^{0\nu}(A, Z)|^2 G^{0\nu}(E_0, Z),$$

$G^{0\nu}(E_0, Z)$ ,  $E_0$  - known phase-space factor and energy release.

If we use a model  $M$  of the calculation of NME,

$$|\langle m \rangle|_M^2(A, Z) = \frac{1}{T_{1/2}^{0\nu}(A, Z) |M_M^{0\nu}(A, Z)|^2 G^{0\nu}(E_0, Z)}.$$

Suppose  $(\beta\beta)_{0\nu}$ -decay of several nuclei is observed.

$|\langle m \rangle|$  cannot depend on parent nucleus  $(A_j, Z_j)$ .

If the light Majorana  $\nu$ -exchange - dominant mechanism of  $(\beta\beta)_{0\nu}$ -decay, **model  $M$  for NME can be correct only if**

$$|\langle m \rangle|_M^2(A_1, Z_1) \simeq |\langle m \rangle|_M^2(A_2, Z_2) = \dots$$

For different models and the same nucleus  $(A, Z)$ ,

$$|\langle m \rangle|_{M_1}^2(A, Z) |M_{M_1}^{0\nu}(A, Z)|^2 = |\langle m \rangle|_{M_2}^2(A, Z) |M_{M_2}^{0\nu}(A, Z)|^2 = \dots,$$

$$|\langle m \rangle|_{M_2}^2(A, Z) = \eta^{M_2; M_1}(A, Z) |\langle m \rangle|_{M_1}^2(A, Z),$$

$$\eta^{M_2; M_1}(A, Z) = \frac{|M_{M_1}^{0\nu}(A, Z)|^2}{|M_{M_2}^{0\nu}(A, Z)|^2}.$$

Nucleus	$\eta^{M_2;M_1}$	$\eta^{M_3;M_1}$	$\eta^{M_2;M_3}$
$^{76}\text{Ge}$	0.37	0.19	1.93
$^{82}\text{Se}$	—	0.38	—
$^{100}\text{Mo}$	—	—	6.56
$^{130}\text{Te}$	0.74	0.10	7.32
$^{136}\text{Xe}$	0.53	0.02	22.42

$M_1$  (SM): E. Caurier et al., 1999;  $M_2$  (QRPA): V. Rodin et al., 2003;  
 $M_3$  (QRPA): O. Civitarese and J. Suhonen, 2003.

The observation of  $(\beta\beta)_{0\nu}$ -decay of at least 3 nuclei would be important for the solution of the problem of NME.

Table 2 suggests:  $^{76}\text{Ge}$ ,  $^{130}\text{Te}$ ,  $^{136}\text{Xe}$  .

If for some model  $M$

$$|\langle m \rangle|_M^2(A_1, Z_1) \simeq |\langle m \rangle|_M^2(A_2, Z_2) = \dots \equiv |\langle m \rangle|_0^2 ,$$

$|\langle m \rangle|_0$  - the true value (most likely).

Strong dependence of NME on  $(A, Z)$  - crucial for the test.

S. M. Bilenky, S.T.P., 2004

Encouraging results on the problem of calculating the NME ( $\xi \lesssim 1.5$ ) have been obtained recently in

V. A. Rodin et al., nucl-th/0503063

$\nu_\odot$ ,  $\Delta m_{\text{atm}}^2$ , CHOOZ Data:

- $\theta_{12} = \theta_\odot \cong \frac{\pi}{6}$ ,  $\theta_{23} = \theta_{\text{atm}} \cong \frac{\pi}{4}$ ,  $\theta_{13} < \frac{\pi}{12}$

$$U = \begin{pmatrix} \frac{\sqrt{3}}{2} & \frac{1}{2} & \epsilon \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & \frac{1}{\sqrt{2}} \\ \frac{1}{2\sqrt{2}} & -\frac{\sqrt{3}}{2\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix} \quad (6)$$

Very different from the CKM-matrix!

- $\cos \theta_{12} \cong \cos(\frac{\pi}{4} - \frac{\pi}{12}) = \frac{1}{\sqrt{2}}(1 + \lambda)$ ,  $\sin \theta_{12} \cong \frac{1}{\sqrt{2}}(1 - \lambda)$ ,
- $\lambda \cong (0.20 - 0.25)$ :  $\theta_\odot + \theta_c = \pi/4$  ?

Natural Possibility:

$$U = U_{\text{lep}}^\dagger(\lambda) U_{\text{bimax}} \quad (7)$$

with

$$U_{\text{bimax}} = \begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} & 0 \\ -\frac{1}{2} & \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{2} & -\frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix} \quad (8)$$

- $U_{\text{lep}}^\dagger(\lambda)$  - from diagonalization of the  $l^-$  mass matrix,
- $U_{\text{bimax}}$  - from diagonalization of the  $\nu$ -mass matrix

Further,  $\Delta m_\odot^2 \ll |\Delta m_{\text{atm}}^2|$ .

- $U_{\text{bimax}}$  can be associated with a symmetry:

$$L' = L_e - L_\mu - L_\tau$$

S.T.P., 1982

This symmetry cannot be exact.

# The Case of CP Nonconservation

$$M_{\text{lep}} = U_L^\dagger m_{\text{lep}}^{\text{diag}} U_R, \quad M_\nu = U_\nu^T m_\nu^{\text{diag}} U_\nu; \quad U = e^{i\Phi} P \tilde{U} Q$$

$$U_{\text{PMNS}} = U_L^\dagger U_\nu = \tilde{U}_{\text{lep}}^\dagger P_\nu \tilde{U}_\nu Q_\nu$$

- $\tilde{U}_{\text{lep}}, \tilde{U}_\nu$  - CKM-like: (3+3) angles, (1+1) CPVP
- $P_\nu = \text{diag}(1, e^{i\phi}, e^{i\omega}), Q_\nu = \text{diag}(1, e^{i\rho}, e^{i\sigma})$ : 4 CPVP

$U_{\text{PMNS}}$ : 3 angles, 3 CPVP

$\tilde{U}_{\text{lep}}^\dagger P_\nu \tilde{U}_\nu Q_\nu$ : 6 angles, 6 CPVP; textures, symmetries

Suppose  $\tilde{U}_\nu$  - bimaximal (real) and arises from

$$M_\nu = \frac{m}{\sqrt{2}} \begin{pmatrix} 0 & e^{-i\alpha'} & e^{-i\beta'} \\ e^{-i\alpha'} & 0 & \epsilon e^{-i\gamma'} \\ e^{-i\beta'} & \epsilon e^{-i\gamma'} & 0 \end{pmatrix},$$

$\alpha', \beta', \gamma'$  - phases,  $\epsilon \ll 1$

$\Delta m_{\text{atm}}^2 \cong m^2, \Delta m_{\odot}^2 \cong \sqrt{2}\epsilon \Delta m_{\text{atm}}^2, \epsilon \sim 0.025$ , IH  $\nu$ - masses

In the limit  $\epsilon = 0$  and  $U_{\text{lep}} = 1$ ,

$L' = L_e - L_\mu - L_\tau$  is conserved.

For  $U_{\text{lep}} \neq 1$ ,  $(\alpha' - \gamma'), (\beta' - \gamma')$  physical CPVP,

$$Q_\nu = 1, \quad P_\nu = \text{diag}(1, e^{i(\beta' - \gamma')}, e^{i(\alpha' - \gamma')})$$

$U_{\text{PMNS}} = \tilde{U}_{\text{lep}}^\dagger P_\nu U_{\text{bimax}}$ : 3 angles, 3 CPVP

P. Frampton, S.T.P., W. Rodejohann, 2004

## Rephasing Invariants Associated with CPVP

Dirac phase  $\delta$ :

$$J_{CP} = \text{Im} \{ U_{e1} U_{\mu 2} U_{e2}^* U_{\mu 1}^* \} .$$

C. Jarlskog, 1985

CP-, T- violation effects in neutrino oscillations

P. Krastev, S.T.P., 1988

Majorana phases  $\alpha_{21}, \alpha_{31}$ :

$$S_1 = \text{Im} \{ U_{e1} U_{e3}^* \}, \quad S_2 = \text{Im} \{ U_{e2} U_{e3}^* \} \quad (\text{not unique})$$

J.F. Nieves and P. Pal, 1987, 2001

G.C. Branco et al., 1986

J.A. Aguilar-Saavedra and G.C. Branco, 2000

$S_1, S_2$  appear in  $|\langle m \rangle|$  in  $(\beta\beta)_{0\nu}$ -decay.

In general,  $J_{CP}, S_1$  and  $S_2$  are independent.

However, for, e.g., all  $\lambda_{ij} \lesssim \lambda$  small, in the model we are considering and to leading order in  $\lambda$ ,

$$J_{CP} \simeq \frac{S_1}{2\sqrt{2}} \simeq \frac{S_2}{2\sqrt{2}} ,$$

and

$$|\langle m \rangle| \cong \sqrt{|\Delta m_{\text{atm}}^2|} |\cos 2\theta_{\odot} + i 8 J_{CP}| .$$

P. Frampton, S.T.P., W. Rodejohann, 2004



## Conclusions

$(\beta\beta)_{0\nu}$ -decay experiments have remarkable physics potential:

- Can establish the Majorana nature of  $\nu_j$
- Can provide unique information on the  $\nu$  mass spectrum
- Can provide unique information on the absolute scale of  $\nu$  masses
- Can provide information on the Majorana CPV phases

The knowledge of the values of the relevant  $(\beta\beta)_{0\nu}$ -decay NME with a sufficiently small uncertainty is crucial for obtaining quantitative information on the neutrino mass and mixing parameters from a measurement of  $\Gamma(\beta\beta)_{0\nu}$ .